

Available online at www.sciencedirect.com



Journal of Molecular Spectroscopy 217 (2003) 195-211

Journal of MOLECULAR SPECTROSCOPY

www.elsevier.com/locate/jms

(000) and (010) States of $H_2^{18}O$: analysis of rotational transitions in hot emission spectrum in the 400–850 cm⁻¹ region

S.N. Mikhailenko,^a Vl.G. Tyuterev,^{b,*} and G. Mellau^c

^a Institute of Atmospheric Optics, Russian Academy of Sciences, 1, av. Akademicheskii, 634055, Tomsk, Russia

^b Groupe de Spectrométrie Moléculaire et Atmosphérique, UMR CNRS 6089, Université de Reims, Faculté des Sciences, Moulin de la Housse, BP 1039,

F-51687 Reims Cedex 2, France

^c Justus-Liebig-Universitaet Giessen, Heinrich-Buff-Ring 58, D-35392 Giessen, Germany

Received 12 July 2002; in revised form 6 September 2002

Abstract

The far-infrared emission spectrum of water vapour have been recorded at temperature 1370 K in the range 400–850 cm⁻¹ at resolutions (FWHM) 0.0055 cm⁻¹. The measurements have been done in a fused quartz cell with an effective length of hot gas of about 60 cm. Among about 200 lines assigned to the H₂¹⁸O isotopomer 106 lines have not been reported in previous studies. These newly measured lines correspond to higher values of rotational quantum numbers compared to previous determinations: $J_{max} = 18$ and $K_{a(max)} = 17$ for the (000) \leftarrow (000) band and $J_{max} = 15$ and $K_{a(max)} = 11$ for the (010) \leftarrow (010) band. Estimated accuracy of line position measurements is 0.0005 cm⁻¹. To our knowledge rotational transitions within an excited bending state have been measured for the first time for H₂¹⁸O. An extended set of experimental rotational energy levels for (000) and (010) vibration states including all previously available data has been determined. A data reduction has been done using the generating function model. The root mean square (RMS) deviation between observed and calculated values is 0.0009 cm⁻¹ for 305 rotational levels of the (000) state and 0.0007 cm⁻¹ for 198 rotational levels of the (010) state. A comparison of the observed data with the best available predictions from the molecular electronic potential energy surface [J. Chem. Phys. 106 (1997) 4618] and isotopic effects for rotational levels of (000) and (010) states are discussed.

© 2003 Elsevier Science (USA). All rights reserved.

Keywords: Far-infrared spectroscopy; Hot water; H₂¹⁸O; Emission spectra; Non-rigid molecules; Centrifugal distortion; Rotational levels; Isotopic effects

1. Introduction

A detailed knowledge of line positions and energy levels of the water molecule and of its isotopic species is known to be important for various atmospheric and astrophysical applications of the spectroscopy. On the other hand spectroscopy of the water molecule stimulate theoretical studies of non-rigid molecules with extremely strong centrifugal distortion. The calculation of water rotational energies, especially for high K_a quantum number, is commonly considered a touchstone for many theoretical models. The aim of this study was to extend experimental information on rotational energy levels of

* Corresponding author. Fax: +33-3-26-91-31-47.

 $H_2^{18}O$ from an analysis of emission spectra of heated water vapour in the 400–850 cm⁻¹ region and to study isotopic effects in observed and calculated levels.

The H₂¹⁸O is the second most abundant isotopomer of water after H₂¹⁶O and has been a subject of rather intensive studies. Rotational energies of the ground and first excited vibrational states of H₂¹⁸O has been determined by Fraley et al. [1] and Williamson et al. [2] through the analysis of spectra in the range of fundamental bands occurring in the near infrared at 5–7.5 µm and at 2.5–3.0 µm region. Few transitions in the v_2 band have been measured by Johns and McKeller [3] using the technique of intracavity CO laser Stark spectroscopy. The v_2 fundamental band has been studied more extensively at high resolution by Guelachvili [4], Camy-Peyret et al. [5], and Toth [6]. Pure rotational transitions

E-mail address: vladimir.tyuterev@univ-reims.fr (V.G. Tyuterev).

in microwave spectra recorded by Steenbeckeliers and Bellet [7] and De Lucia et al. [8] and in far-infrared spectra recorded by Johns [9] in the 60–280 cm⁻¹ region and by Kauppinen and Kyro [10] in the 67–695 cm⁻¹ region have contributed to an accurate determination of rotational energy levels of H₂¹⁸O with relatively lowrotational J, K_a quantum numbers. New measurements of (000) \leftarrow (000) rotational transitions have been recently reported by Toth [11] in 595–943 cm⁻¹ region and by Matsushima et al. [12] in 18–172 cm⁻¹ region. Spectra of H₂¹⁸O have been studied further in the region of the 1st triad [13–16] and the 2nd triad [17,18] and also in a higher wavenumber range [19–20].

Note that all experimentally available rotational levels [11] and almost all rovibrational levels of $H_2^{18}O$ were determined from analysis of room-temperature spectra and thus correspond to relatively moderate values of rotational quantum numbers J, K_a . The only hot $H_2^{18}O$ band analysed so far is the $(021) \leftarrow (010)$ band reported by Chevillard et al. [21]. Also some $H_2^{18}O v_3$ line positions at 1200 K with natural isotopic abundance have been measured by Pine et al. [22]. The situation is quite different from that of the major isotopomer $H_2^{16}O$ for which a lot of information on high J, K_a energy levels has been deduced from hot water spectra (sf. for example [22–28] and references therein).

In this paper we report a first observation of rotational transitions $(010) \leftarrow (010)$ in the excited bending vibration state of H₂¹⁸O in water emission spectra recorded at 1370 K. Also new rotational transitions in the ground vibration state (000) have been assigned that allows to improve an accuracy of the determination of highly excited rotational levels and to extend considerably the range of experimentally accessible rotational quantum numbers for H₂¹⁸O.

The data reduction has been done using the generating-function model [29–31] of the effective rotational Hamiltonian because the usual polynomial model for centrifugal distortion terms [32] employed in all previous analysis of H₂¹⁸O spectra [1–20] does not provide a satisfactory accuracy for our data. Comparisons of our experimental data and calculations with global predictions for (000) and (010) levels performed by Partridge and Schwenke [33] from an electronic potential energy surface, as well as isotopic shift variation versus J and K_a are discussed.

2. Experimental spectra analysis

The far-infrared emission spectrum of heated water vapour (with natural isotopic abundance) has been recorded at resolution 0.0055 cm^{-1} at Giessen on the Brucker IFS 120 HR Fourier transform spectrometer [34] which had been previously used to record room-temperature spectra of $D_2^{16}O$ [35] and $H_2^{16}O$ [28,36].

The measurements for the present work were done in a fused quartz cell with an effective length of hot gas of about 60 cm. The gas was at temperature of 1370 K although a small segment near the windows was at much lower temperature. A description of the cell and other experimental details have been given in papers on HCN emission spectra experiments [37,38] in similar experimental conditions. Table 1 lists the conditions used in the recording of the emission spectrum.

Because the water sample had a natural isotopic abundance, most of ~13000 emission water lines recorded in the 400–850 cm⁻¹ region in the present study belong to the H₂¹⁶O isotopomer. These lines together with previously available data had been used in [28] to improve an accuracy of a determination of some H₂¹⁶O rovibrational states. About 200 lines have been assigned to $(000) \leftarrow (000)$ and $(010) \leftarrow (010)$ bands of the H₂¹⁸O isotopomer.

Low-J, K_a transitions could be readily assigned using available database compilations [39], and associated line positions are in a good agreement with roomtemperature measurements [10,11]. An assignment of higher J, K_a transitions corresponding to weak observed lines requires extrapolations based on theoretical calculations. Because the standard polynomial model of the effective Hamiltonian in the case of watertype non-rigid molecules has poor convergence and extrapolation properties, we have used the generating function model [29-31] outlined in the following section, in a similar way as in our previous work on $H_2^{16}O$ spectra [28,36]. The assignment of transitions has been done using step-by-step iterative extrapolations with increasing J and K_a quantum numbers. The calibration of line positions has been achieved using available literature data [10,11] as well as our previous measurements of H₂¹⁶O lines [28].

In Table 2 we give new line positions of $H_2^{18}O$ (for transitions which had not been observed in previously reported room-temperature spectra) together with their rotational assignments. The second column gives estimations (in 10^{-4} cm⁻¹) of absolute uncertainties in experimental line positions.

Table 1

Laboratory measurement conditions for emission FTIR spectra of water vapour

| Spectrometer | Bruker IFS 120 HR |
|----------------------|--------------------------------------|
| Gas temperature | 1370 K |
| Total water pressure | $1.6 	imes 10^3 \mathrm{Pa}$ |
| Length of the cell | 1 m |
| Diameter of the cell | 6 cm |
| Length of hot zone | 60 cm |
| Detector | Ge:Cu at 4K |
| Filter | $400-900{\rm cm^{-1}}$ at $4{\rm K}$ |
| Window | KBr |
| Resolution | $0.0055{ m cm}^{-1}$ |
| | |

Table 2 New ${\rm H_2}^{18}O$ lines observed in emission spectrum

| Observed | σ | Observed | 0 – C | Upper state | | | | Lower state | | | | |
|-------------|-------------|-----------|-------|---------------|----|----|----|---------------|----|----|----------------|--|
| position | | emittance | | $v_1 v_2 v_3$ | J | Ka | Kc | $v_1 v_2 v_3$ | J | Ka | K _c | |
| (a) (000) ← | (000) trans | itions | | | | | | | | | | |
| 414.6743 | 5 | 0.043 | -3 | 000 | 9 | 8 | 1 | 000 | 8 | 7 | 2 | |
| 419.0664 | 5 | 0.020 | 10 | 000 | 10 | 7 | 3 | 000 | 9 | 6 | 4 | |
| 418.8777 | 5 | 0.026 | -2 | 000 | 10 | 7 | 4 | 000 | 9 | 6 | 3 | |
| 439.2281 | 5 | 0.032 | -5 | 000 | 10 | 8 | 2 | 000 | 9 | 7 | 3 | |
| 439.2219 | 5 | 0.060 | -3 | 000 | 10 | 8 | 3 | 000 | 9 | 7 | 2 | |
| 456.6028 | 5 | 0.119 | -7 | 000 | 10 | 9 | 2 | 000 | 9 | 8 | 1 | |
| 471.2257 | 5 | 0.146 | -8 | 000 | 10 | 10 | 1 | 000 | 9 | 9 | 0 | |
| 547.7524 | 5 | 0.023 | -3 | 000 | 11 | 4 | 8 | 000 | 10 | 1 | 9 | |
| 416.4078 | 5 | 0.032 | 12 | 000 | 11 | 5 | 6 | 000 | 10 | 4 | 7 | |
| 543.2684 | 5 | 0.021 | 4 | 000 | 11 | 5 | 7 | 000 | 10 | 2 | 8 | |
| 442.8515 | 5 | 0.070 | 0 | 000 | 11 | 7 | 4 | 000 | 10 | 6 | 5 | |
| 463.4088 | 5 | 0.051 | -4 | 000 | 11 | 8 | 3 | 000 | 10 | 7 | 4 | |
| 457.5228 | 5 | 0.032 | -4 | 000 | 12 | 5 | 7 | 000 | 11 | 4 | 8 | |
| 466.4398 | 5 | 0.025 | 0 | 000 | 12 | 7 | 5 | 000 | 11 | 6 | 6 | |
| 505.5149 | 20 | 0.042 | 4 | 000 | 12 | 9 | 3 | 000 | 11 | 8 | 4 | |
| 505.5092 | 5 | 0.120 | 2 | 000 | 12 | 9 | 4 | 000 | 11 | 8 | 3 | |
| 534.0296 | 5 | 0.153 | 11 | 000 | 12 | 11 | 2 | 000 | 11 | 10 | 1 | |
| 544.3549 | 5 | 0.169 | -5 | 000 | 12 | 12 | 1 | 000 | 11 | 11 | 0 | |
| 623.2937 | 5 | 0.021 | 5 | 000 | 13 | 5 | 9 | 000 | 12 | 2 | 10 | |
| 441.8879 | 20 | 0.032 | -4 | 000 | 13 | 6 | 8 | 000 | 12 | 5 | 7 | |
| 490.1788 | 5 | 0.128 | -2 | 000 | 13 | 7 | 6 | 000 | 12 | 6 | 7 | |
| 484.9149 | 5 | 0.032 | -1 | 000 | 13 | 7 | 7 | 000 | 12 | 6 | 6 | |
| 510.0354 | 5 | 0.035 | 0 | 000 | 13 | 8 | 6 | 000 | 12 | 7 | 5 | |
| 529.3096 | 5 | 0.035 | -23 | 000 | 13 | 9 | 4 | 000 | 12 | 8 | 5 | |
| 529.2881 | 5 | 0.021 | 0 | 000 | 13 | 9 | 5 | 000 | 12 | 8 | 4 | |
| 545.4061 | 5 | 0.141 | 0 | 000 | 13 | 10 | 3 | 000 | 12 | 9 | 4 | |
| 558.8406 | 5 | 0.122 | 5 | 000 | 13 | 11 | 2 | 000 | 12 | 10 | 3 | |
| 569.7219 | 20 | 0.298 | -8 | 000 | 13 | 12 | 1 | 000 | 12 | 11 | 2 | |
| 578.0844 | 5 | 0.108 | 2 | 000 | 13 | 13 | 0 | 000 | 12 | 12 | 1 | |
| 559.3041 | 10 | 0.022 | 14 | 000 | 14 | 5 | 9 | 000 | 13 | 4 | 10 | |
| 668.8314 | 5 | 0.021 | 14 | 000 | 14 | 5 | 10 | 000 | 13 | 2 | 11 | |
| 513.9257 | 5 | 0.025 | -15 | 000 | 14 | 6 | 8 | 000 | 13 | 5 | 9 | |
| 664.4677 | 5 | 0.021 | 17 | 000 | 14 | 6 | 9 | 000 | 13 | 3 | 10 | |
| 514.7348 | 5 | 0.025 | -4 | 000 | 14 | 7 | 7 | 000 | 13 | 6 | 8 | |
| 719.3791 | 10 | 0.021 | 12 | 000 | 14 | 7 | 8 | 000 | 13 | 4 | 9 | |
| 502.7224 | 5 | 0.032 | 2 | 000 | 14 | 7 | 8 | 000 | 13 | 6 | 7 | |
| 533.3844 | 5 | 0.027 | -7 | 000 | 14 | 8 | 6 | 000 | 13 | 7 | 7 | |
| 532.1362 | 5 | 0.057 | 0 | 000 | 14 | 8 | 7 | 000 | 13 | 7 | 6 | |
| 552.6054 | 5 | 0.023 | -14 | 000 | 14 | 9 | 5 | 000 | 13 | 8 | 6 | |
| 552.5208 | 5 | 0.075 | 2 | 000 | 14 | 9 | 6 | 000 | 13 | 8 | 5 | |
| 569.2275 | 5 | 0.088 | 23 | 000 | 14 | 10 | 5 | 000 | 13 | 9 | 4 | |
| 583.1719 | 5 | 0.108 | 6 | 000 | 14 | 11 | 4 | 000 | 13 | 10 | 3 | |
| 594.5862 | 5 | 0.097 | 6 | 000 | 14 | 12 | 3 | 000 | 13 | 11 | 2 | |
| 603.5554 | 5 | 0.095 | 14 | 000 | 14 | 13 | 2 | 000 | 13 | 12 | 1 | |
| 610.0751 | 5 | 0.092 | 3 | 000 | 14 | 14 | 1 | 000 | 13 | 13 | 0 | |
| 694.0695 | 5 | 0.021 | -23 | 000 | 15 | 4 | 11 | 000 | 14 | 3 | 12 | |
| 617.3893 | 15 | 0.021 | 0 | 000 | 15 | 5 | 10 | 000 | 14 | 4 | 11 | |
| 555.1027 | 5 | 0.029 | 1 | 000 | 15 | 6 | 9 | 000 | 14 | 5 | 10 | |
| 699.2696 | 20 | 0.030 | 20 | 000 | 15 | 6 | 10 | 000 | 14 | 3 | 11 | |
| 452.9466 | 5 | 0.032 | -9 | 000 | 15 | 6 | 10 | 000 | 14 | 5 | 9 | |
| 541.1532 | 5 | 0.043 | -19 | 000 | 15 | 7 | 8 | 000 | 14 | 6 | 9 | |
| 516.7371 | 5 | 0.032 | 18 | 000 | 15 | 7 | 9 | 000 | 14 | 6 | 8 | |
| 556.0914 | 5 | 0.052 | -5 | 000 | 15 | 8 | 7 | 000 | 14 | 7 | 8 | |
| 552.9179 | 5 | 0.023 | 8 | 000 | 15 | 8 | 8 | 000 | 14 | 7 | 7 | |
| 575.3734 | 5 | 0.063 | 3 | 000 | 15 | 9 | 6 | 000 | 14 | 8 | 7 | |
| 575.1055 | 5 | 0.021 | 14 | 000 | 15 | 9 | 7 | 000 | 14 | 8 | 6 | |
| 592.5294 | 5 | 0.068 | 7 | 000 | 15 | 10 | 5 | 000 | 14 | 9 | 6 | |
| 592.5136 | 5 | 0.024 | 13 | 000 | 15 | 10 | 6 | 000 | 14 | 9 | 5 | |
| 606.9880 | 5 | 0.105 | 4 | 000 | 15 | 11 | 4 | 000 | 14 | 10 | 5 | |
| 618 9167 | 5 | 0.097 | -3 | 000 | 15 | 12 | 3 | 000 | 14 | 11 | 4 | |

Table 2 (continued)

| Observed | σ | Observed | O-C | Upper state | | | | Lower state | | | | | |
|---------------------------------|-----------|-----------|----------|---------------|----|--------|----------------|---------------|----------|----------------|----------------|--|--|
| position | | emittance | | $v_1 v_2 v_3$ | J | Ka | K _c | $v_1 v_2 v_3$ | J | K _a | K _c | | |
| 628.4488 | 5 | 0.076 | 0 | 000 | 15 | 13 | 2 | 000 | 14 | 12 | 3 | | |
| 635.6308 | 30 | 0.050 | -31 | 000 | 15 | 14 | 1 | 000 | 14 | 13 | 2 | | |
| 640.4363 | 5 | 0.070 | -12 | 000 | 15 | 15 | 0 | 000 | 14 | 14 | 1 | | |
| 765.2243 | 15 | 0.021 | -44 | 000 | 16 | 5 | 12 | 000 | 15 | 2 | 13 | | |
| 738.5161 | 20 | 0.021 | -46 | 000 | 16 | 6 | 11 | 000 | 15 | 3 | 12 | | |
| 454.5252 | 5 | 0.032 | -32 | 000 | 16 | 6 | 11 | 000 | 15 | 5 | 10 | | |
| 570.8129 | 5 | 0.021 | -32 | 000 | 16 | / | 9 | 000 | 15 | 6 | 10 | | |
| 520.2050 | 5 | 0.052 | -0 | 000 | 16 | 8 | 10 Q | 000 | 15 | 0 | 9 | | |
| 597 6253 | 5 | 0.032 | -3 | 000 | 16 | 9 | 7 | 000 | 15 | 8 | 8 | | |
| 596.8841 | 10 | 0.060 | 6 | 000 | 16 | 9 | 8 | 000 | 15 | 8 | 7 | | |
| 615.2653 | 5 | 0.026 | -24 | 000 | 16 | 10 | 6 | 000 | 15 | 9 | 7 | | |
| 615.2147 | 30 | 0.143 | 16 | 000 | 16 | 10 | 7 | 000 | 15 | 9 | 6 | | |
| 630.2567 | 15 | 0.044 | 3 | 000 | 16 | 11 | 5 | 000 | 15 | 10 | 6 | | |
| 630.2536 | 5 | 0.152 | 3 | 000 | 16 | 11 | 6 | 000 | 15 | 10 | 5 | | |
| 642.6899 | 5 | 0.141 | -25 | 000 | 16 | 12 | 5 | 000 | 15 | 11 | 4 | | |
| 652.7515 | 5 | 0.100 | -19 | 000 | 16 | 13 | 4 | 000 | 15 | 12 | 3 | | |
| 660.5391 | 5 | 0.075 | 12 | 000 | 16 | 14 | 3 | 000 | 15 | 13 | 2 | | |
| 666.0742 | 5 | 0.048 | 73 | 000 | 16 | 15 | 2 | 000 | 15 | 14 | 1 | | |
| 669.2753 | 5 | 0.033 | 17 | 000 | 16 | 16 | 1 | 000 | 15 | 15 | 0 | | |
| 806.2560 | 5 | 0.021 | 269 | 000 | 17 | 4 | 13 | 000 | 16 | 3 | 14 | | |
| /3/.8213 | 5 15 | 0.034 | -9 10 | 000 | 17 | 5 | 12 | 000 | 16 16 | 4 | 13 | | |
| 605 1724 | 15 | 0.021 | 19 | 000 | 17 | 07 | 10 | 000 | 16 | 5 | 12 | | |
| 602 7011 | 5 | 0.021 | 40 | 000 | 17 | 8 | 9 | 000 | 16 | 0 | 10 | | |
| 587 6652 | 5 | 0.020 | -50 | 000 | 17 | 8 | 10 | 000 | 16 | 7 | 9 | | |
| 619.4510 | 5 | 0.021 | -34 | 000 | 17 | 9 | 8 | 000 | 16 | 8 | 9 | | |
| 637.4225 | 5 | 0.028 | -2 | 000 | 17 | 10 | 7 | 000 | 16 | 9 | 8 | | |
| 637.2627 | 5 | 0.021 | 12 | 000 | 17 | 10 | 8 | 000 | 16 | 9 | 7 | | |
| 652.9476 | 5 | 0.032 | -4 | 000 | 17 | 11 | 6 | 000 | 16 | 10 | 7 | | |
| 652.9369 | 5 | 0.021 | -3 | 000 | 17 | 11 | 7 | 000 | 16 | 10 | 6 | | |
| 665.8872 | 5 | 0.063 | -22 | 000 | 17 | 12 | 5 | 000 | 16 | 11 | 6 | | |
| 676.4552 | 30 | 0.080 | 12 | 000 | 17 | 13 | 4 | 000 | 16 | 12 | 5 | | |
| 684.7872 | 5 | 0.093 | 25 | 000 | 17 | 14 | 3 | 000 | 16 | 13 | 4 | | |
| 690.9328 | 5 | 0.035 | -162 | 000 | 17 | 15 | 2 | 000 | 16 | 14 | 3 | | |
| 694.9360 | 5 | 0.041 | -56 | 000 | 17 | 16 | 1 | 000 | 16 | 15 | 2 | | |
| 522 6222 | 20 | 0.038 | -10 | 000 | 17 | 1/ | 12 | 000 | 10 | 10 | 1 11 | | |
| 599 7988 | 20 | 0.021 | _14 | 000 | 18 | 8 | 12 | 000 | 17 | 0 | 10 | | |
| 636 8756 | 5 | 0.021 | -60 | 000 | 18 | 9 | 10 | 000 | 17 | 8 | 9 | | |
| 675.0376 | 5 | 0.026 | -3 | 000 | 18 | 11 | 7 | 000 | 17 | 10 | 8 | | |
| 675.0044 | 5 | 0.038 | 0 | 000 | 18 | 11 | 8 | 000 | 17 | 10 | 7 | | |
| 688.4847 | 5 | 0.071 | -3 | 000 | 18 | 12 | 7 | 000 | 17 | 11 | 6 | | |
| 699.5360 | 20 | 0.023 | -5 | 000 | 18 | 13 | 6 | 000 | 17 | 12 | 5 | | |
| 708.3811 | 15 | 0.021 | 98 | 000 | 18 | 14 | 5 | 000 | 17 | 13 | 4 | | |
| 715.0971 | 15 | 0.021 | -13 | 000 | 18 | 15 | 4 | 000 | 17 | 14 | 3 | | |
| (b) $(0 \ 1 \ 0) \leftarrow (0$ | 010) tran | sitions | | | | | | | | | | | |
| 416.4196 | 5 | 0.032 | -2 | 010 | 8 | 8 | 1 | 010 | 7 | 7 | 0 | | |
| 441.4611 | 5 | 0.021 | -6 | 010 | 9 | 8 | 1 | 010 | 8 | 7 | 2 | | |
| 458.3014 | 5 | 0.032 | 3 | 010 | 9 | 9 | 0 | 010 | 8 | 8 | 1 | | |
| 445.4744 | 5 | 0.032 | 3 | 010 | 10 | 7 | 3 | 010 | 9 | 6 | 4 | | |
| 443.3029 466 1826 | 5 | 0.032 | 1 _/ | 010 | 10 | / Q | 4 | 010 | 9 | 0 7 | 3 | | |
| 400.1620 | 5 | 0.025 | -4 1 | 010 | 10 | o Q | 3 2 | 010 | 9 | / & | ے 1 | | |
| 497 5954 | 5 | 0.074 | + 1 | 010 | 10 | 10 | 2 1 | 010 | 9 9 | 0 9 | 0 | | |
| 446.1404 | 15 | 0.032 | -18 | 010 | 11 | 6 | 5 | 010 | 10 | 5 | 6 | | |
| 469.3371 | 5 | 0.032 | 4 | 010 | 11 | 7 | 4 | 010 | 10 | 6 | 5 | | |
| 490.5236 | 5 | 0.027 | 2 | 010 | 11 | 8 | 3 | 010 | 10 | 7 | 4 | | |
| 490.5053 | 5 | 0.032 | -1 | 010 | 11 | 8 | 4 | 010 | 10 | 7 | 3 | | |
| 508.3240 | 5 | 0.026 | 3 | 010 | 11 | 9 | 2 | 010 | 10 | 8 | 3 | | |
| 522.9247 | 5 | 0.035 | 2 | 010 | 11 | 10 | 1 | 010 | 10 | 9 | 2 | | |

Table 2 (continued)

| Observed | σ | Observed | O-C | Upper st | ate | | | Lower state | | | | |
|----------|----|-----------|------|---------------|-----|----|----------------|---------------|----|----|----------------|--|
| position | | emittance | | $v_1 v_2 v_3$ | J | Ka | K _c | $v_1 v_2 v_3$ | J | Ka | K _c | |
| 534.5378 | 10 | 0.032 | 2 | 010 | 11 | 11 | 0 | 010 | 10 | 10 | 1 | |
| 399.3049 | 5 | 0.020 | -32 | 010 | 12 | 5 | 8 | 010 | 11 | 4 | 7 | |
| 458.6207 | 5 | 0.076 | 41 | 010 | 12 | 6 | 7 | 010 | 11 | 5 | 6 | |
| 492.8576 | 20 | 0.026 | 6 | 010 | 12 | 7 | 5 | 010 | 11 | 6 | 6 | |
| 491.6386 | 20 | 0.032 | 17 | 010 | 12 | 7 | 6 | 010 | 11 | 6 | 5 | |
| 514.4060 | 20 | 0.020 | -41 | 010 | 12 | 8 | 4 | 010 | 11 | 7 | 5 | |
| 514.3407 | 20 | 0.020 | 20 | 010 | 12 | 8 | 5 | 010 | 11 | 7 | 4 | |
| 532.7321 | 5 | 0.029 | 5 | 010 | 12 | 9 | 4 | 010 | 11 | 8 | 3 | |
| 547.8407 | 5 | 0.025 | 1 | 010 | 12 | 10 | 3 | 010 | 11 | 9 | 2 | |
| 559.9877 | 5 | 0.023 | 1 | 010 | 12 | 11 | 2 | 010 | 11 | 10 | 1 | |
| 499.1862 | 5 | 0.032 | 24 | 010 | 13 | 6 | 7 | 010 | 12 | 5 | 8 | |
| 516.2172 | 20 | 0.021 | -45 | 010 | 13 | 7 | 6 | 010 | 12 | 6 | 7 | |
| 556.6711 | 5 | 0.032 | -15 | 010 | 13 | 9 | 4 | 010 | 12 | 8 | 5 | |
| 572.3075 | 5 | 0.021 | 73 | 010 | 13 | 10 | 3 | 010 | 12 | 9 | 4 | |
| 584.9628 | 5 | 0.021 | 1 | 010 | 13 | 11 | 2 | 010 | 12 | 10 | 3 | |
| 532.3723 | 5 | 0.032 | 3 | 010 | 14 | 7 | 8 | 010 | 13 | 6 | 7 | |
| 560.0346 | 5 | 0.023 | 7 | 010 | 14 | 8 | 7 | 010 | 13 | 7 | 6 | |
| 580.0362 | 5 | 0.021 | -138 | 010 | 14 | 9 | 6 | 010 | 13 | 8 | 5 | |
| 596.2389 | 5 | 0.021 | -218 | 010 | 14 | 10 | 5 | 010 | 13 | 9 | 4 | |
| 609.4180 | 5 | 0.023 | -70 | 010 | 14 | 11 | 4 | 010 | 13 | 10 | 3 | |
| 583.2269 | 5 | 0.022 | 0 | 010 | 15 | 8 | 7 | 010 | 14 | 7 | 8 | |
| 602.9489 | 5 | 0.021 | 2 | 010 | 15 | 9 | 6 | 010 | 14 | 8 | 7 | |

Note. Observed line positions are given in cm⁻¹. σ , the error estimation of the line experimental line position, in units 10^{-4} cm⁻¹. Observed emittance is in arbitrary units. O - C, deviation between the observed line position and our calculated line position using parameters of Table 4, in units 10^{-4} cm⁻¹.

3. Data reduction: energy levels and hamiltonian parameters for (000) and (010) states

To assure a consistency with calibrations of previous works on water spectra [1-21], and a self-consistency of energy level determinations from various sources of experimental line positions, we have applied the RITZ code written by Tashkun et al. [40] in a similar way as in our recent study on H₂¹⁶O spectra [28]. This used the RITZ combination principle to recover all possible energy levels by simultaneous processing of line positions in various records over an interval which extended beyond that which is the focus of this work. This procedure accounts for our own measured line positions together with complementary literature data to generate a set of energy levels which provide transition wavenumbers in agreement with observations according to a least squares procedure. In cases where a level is involved in many observed transitions the program provides for averaging, which eliminates to a certain extent contributions of line shifts and gives more reliable experimental estimates than a classical combination difference approach. A detailed description of this technique applied to various molecules will be discussed elsewhere [40].

A total of 3937 observed transitions including those of [6,8-12,15,16,19,21] were simultaneously used to obtain a new set of rotational energies for (000) and (010) vibration states of H₂¹⁸O. The results of energy level

determinations achieved with this procedure are presented in Table 3 in comparison with the most complete previous set of [11]. The column headed " N_t " indicates number of observed transitions involved in the determination of each level. A statistical analysis of the obtained data set, performed with the RITZ-program [40], indicates that information from microwave transitions [8] and particularly from very accurate recent measurements of Matsushima et al. [12] in 1–5 THz region contributes significantly to precise determination of low-*J* levels, whereas complementary information from our high-temperature spectra allow to determine higher rotational energies.

It is well known that an accurate calculation of excited rotational states of the water molecule and of related analysis of high-resolution spectra is complicated by an extremely strong centrifugal distortion resulting from the bending-rotational coupling. Because of this coupling the standard power series expansion of the effective rotational Hamiltonian H_{rot} used in previous papers pertaining to the assignment of the H₂¹⁸O spectra has a limited domain of validity. According to estimations [29] of the convergence radius of this expansion, the domain of its applicability is limited by a maximal value $K_{a max}$ which depends on v_2 . For (000) and (010) states of the water molecule the convergence radii of the power series expansion of $H_{\rm rot}$ correspond to $K_{a max} = 11$ and $K_{a max} = 9$, but we have observed transitions with higher values of K_a : up to 17 in the case

| Table 3 |
|---|
| Experimentally determined rotation energies of (000) and (010) vibration states of H ₂ ¹⁸ O |

| J | Ka | Kc | (000) State | - | · / | | (010) State | | | | | |
|---|----|----|------------------|-------------------|-----|-----------|-------------|------------------|-------------------|-------------|-----------|------------|
| | | | $E_{\rm obs}/hc$ | $\delta_{ m obs}$ | Nt | Obs – Cal | Obs - [11] | $E_{\rm obs}/hc$ | $\delta_{ m obs}$ | $N_{\rm t}$ | Obs – Cal | Obs - [11] |
| 1 | | | 2 | 3 | 4 | 5 | 6 | 2 | 3 | 4 | 5 | 6 |
| 0 | 0 | 0 | 0 | | | | | 1588.27563 | .16 | 2 | .00 | 04 |
| 1 | 0 | 1 | 23.75490 | .02 | 21 | .03 | .00 | 1612.04599 | .13 | 4 | .03 | 02 |
| 1 | 1 | 1 | 36.74859 | .03 | 22 | 18 | 06 | 1628.06007 | .11 | 6 | .11 | .01 |
| 1 | 1 | 0 | 42.02341 | .02 | 20 | 07 | 02 | 1633.63470 | .15 | 3 | .10 | 01 |
| 2 | 0 | 2 | 69.92736 | .03 | 26 | 01 | 08 | 1658.33384 | .08 | 7 | 06 | 06 |
| 2 | 1 | 2 | 78.98863 | .02 | 35 | 13 | 02 | 1670.03713 | .11 | 7 | .08 | .02 |
| 2 | 1 | 1 | 94.78855 | .03 | 25 | 04 | 10 | 1686.73330 | .08 | 6 | .08 | 07 |
| 2 | 2 | 1 | 133.47580 | .02 | 30 | 10 | .00 | 1734.21855 | .09 | 8 | .13 | .03 |
| 2 | 2 | 0 | 134.78304 | .03 | 29 | 17 | 08 | 1735.43765 | .09 | 6 | 04 | 09 |
| 3 | 0 | 3 | 136.33667 | .02 | 42 | .07 | .00 | 1725.01686 | .07 | 8 | 05 | 03 |
| 3 | 1 | 3 | 141.56799 | .03 | 36 | 16 | 07 | 1732.26271 | .07 | 8 | .00 | 06 |
| 3 | 1 | 2 | 172.88291 | .02 | 39 | .14 | .00 | 1765.39797 | .08 | 8 | .09 | 04 |
| 3 | 2 | 2 | 204.75584 | .03 | 34 | 11 | 07 | 1805.57662 | .06 | 8 | 02 | 09 |
| 3 | 2 | 1 | 210.79925 | .02 | 45 | 02 | 03 | 1811.29059 | .08 | 8 | .00 | 03 |
| 3 | 3 | 1 | 282.09453 | .03 | 28 | .05 | 05 | 1897.45280 | .05 | 9 | 25 | 02 |
| 3 | 3 | 0 | 282.30712 | .02 | 31 | .12 | .03 | 1897.62728 | .07 | 8 | 17 | .04 |
| 4 | 0 | 4 | 221.23389 | .02 | 36 | 09 | 12 | 1810.18727 | .09 | 8 | 10 | 05 |
| 4 | 1 | 4 | 223.82851 | .02 | 41 | 09 | 05 | 1814.08713 | .05 | 8 | .04 | 03 |
| 4 | 1 | 3 | 274.80317 | .02 | 33 | .19 | 04 | 1868.25382 | .03 | 9 | .08 | 03 |
| 4 | 2 | 3 | 298.62016 | .02 | 46 | .02 | 02 | 1899.60893 | .04 | 11 | .05 | 01 |
| 4 | 2 | 2 | 314.45937 | .02 | 42 | 03 | 09 | 1914.88087 | .06 | 8 | 11 | 04 |
| 4 | 3 | 2 | 379.29164 | .02 | 46 | .07 | .02 | 1994.70331 | .06 | 11 | 18 | .00 |
| 4 | 3 | 1 | 380.70248 | .03 | 31 | .02 | 04 | 1995.87007 | .03 | 7 | 31 | 06 |
| 4 | 4 | 1 | 482.64365 | .05 | 24 | .18 | .08 | 2117.00851 | .06 | 9 | .03 | .02 |
| 4 | 4 | 0 | 482.67261 | .03 | 28 | .07 | 01 | 2117.02965 | .09 | 6 | 04 | 06 |
| 5 | 0 | 5 | 324.04672 | .02 | 39 | 04 | 03 | 1913.02263 | .04 | 9 | 01 | 01 |
| 5 | 1 | 5 | 325.21569 | .02 | 34 | 10 | 05 | 1914.93344 | .05 | 9 | .01 | 08 |
| 5 | 1 | 4 | 398.36049 | .02 | 44 | .21 | 03 | 1993.27685 | .08 | 10 | .02 | 04 |
| 5 | 2 | 4 | 414.16813 | .02 | 38 | .02 | 06 | 2015.45390 | .07 | 9 | .05 | 06 |
| 5 | 2 | 3 | 445.15853 | .02 | 47 | .05 | 05 | 2045.92261 | .08 | 10 | 20 | 06 |
| 5 | 3 | 3 | 500.59625 | .02 | 36 | .04 | 02 | 2116.13852 | .07 | 10 | 25 | 06 |
| 5 | 3 | 2 | 505.72883 | .02 | 48 | .04 | 01 | 2120.45465 | .08 | 10 | 24 | .00 |
| 5 | 4 | 2 | 604.54420 | .02 | 35 | .01 | 02 | 2238.97487 | .05 | 9 | .09 | 08 |
| 5 | 4 | 1 | 604.79296 | .02 | 37 | .08 | .02 | 2239.15709 | .05 | 10 | .19 | .04 |
| 5 | 5 | 1 | 733.67938 | .03 | 12 | 17 | 01 | 2390.35364 | .03 | 4 | .36 | .06 |
| 5 | 5 | 0 | 733.68318 | .02 | 21 | 07 | .08 | 2390.35607 | .06 | 5 | .37 | 01 |
| 6 | 0 | 6 | 444.84621 | .02 | 28 | 02 | 02 | 2033.48861 | .06 | 8 | 06 | 09 |
| 6 | 1 | 6 | 445.34619 | .02 | 39 | 07 | 04 | 2034.37686 | .05 | 10 | .02 | 05 |
| 6 | 1 | 5 | 541.18006 | .02 | 31 | .12 | 11 | 2138.08857 | .05 | 8 | .04 | 02 |
| 6 | 2 | 5 | 550.45078 | .02 | 38 | .06 | 07 | 2152.18203 | .06 | 11 | .10 | 09 |
| 6 | 2 | 4 | 601.23775 | .02 | 36 | .02 | 11 | 2202.99269 | .03 | 8 | 22 | 06 |
| 6 | 3 | 4 | 645.38254 | .02 | 42 | .00 | 09 | 2261.21754 | .06 | 11 | 11 | .01 |
| 6 | 3 | 3 | 658.61006 | .02 | 39 | 04 | 04 | 2272.64610 | .09 | 8 | 29 | 04 |
| 6 | 4 | 3 | 751.03306 | .02 | 46 | .11 | .05 | 2385.51357 | .06 | 11 | .17 | .00 |
| 6 | 4 | 2 | 752.18744 | .03 | 36 | 01 | 08 | 2386.36584 | .06 | 9 | .08 | 03 |
| 6 | 5 | 2 | 880.07644 | .02 | 27 | 08 | .08 | 2536.90968 | .05 | 7 | .81 | .01 |
| 6 | 5 | 1 | 880.11451 | .03 | 30 | 30 | 12 | 2536.93482 | .03 | 7 | .78 | 07 |
| 6 | 6 | 1 | 1033.19417 | .03 | 15 | 20 | .03 | 2714.44446 | .09 | 5 | .11 | .10 |
| 6 | 6 | 0 | 1033.19450 | .03 | 6 | 33 | 06 | 2714.44446 | .04 | 2 | 17 | .00 |
| 7 | 0 | 7 | 583.77783 | .02 | 34 | 04 | .03 | 2171.74067 | .07 | 8 | .05 | 07 |
| 7 | 1 | 7 | 583.98642 | .02 | 26 | 08 | 11 | 2172.14323 | .07 | 7 | .12 | 01 |
| 7 | 1 | 6 | 701.69426 | .02 | 39 | .13 | 07 | 2300.77431 | .05 | 10 | .00 | .04 |
| 7 | 2 | 6 | 706.59781 | .02 | 37 | .16 | 02 | 2308.90599 | .08 | 7 | .15 | 07 |
| 7 | 2 | 5 | 780.45278 | .02 | 41 | 06 | 08 | 2384.04190 | .04 | 9 | 23 | 09 |
| 7 | 3 | 5 | 812.76174 | .03 | 33 | .15 | 02 | 2429.13499 | .03 | 8 | .00 | 05 |
| 7 | 3 | 4 | 839.54949 | .02 | 40 | 05 | 05 | 2453.05766 | .05 | 9 | 07 | .03 |
| 7 | 4 | 4 | 921.89593 | .03 | 30 | .11 | .03 | 2556.46209 | .03 | 8 | .06 | 08 |
| 7 | 4 | 3 | 925.69992 | .02 | 36 | .22 | .09 | 2559.31394 | .04 | 8 | .07 | 05 |
| 7 | 5 | 3 | 1050.99035 | .03 | 28 | 03 | .05 | 2707.92161 | .03 | 6 | .52 | 12 |

Table 3 (continued)

| $ \begin{array}{ c c c c c c c c c c c c c c c c c c c$ | Obs – Cal | 01 (111) |
|---|-----------|-------------|
| $ \begin{array}{ c c c c c c c c c c c c c c c c c c c$ | | Obs - [11] |
| $ \begin{array}{cccccccccccccccccccccccccccccccccccc$ | 5 | 6 |
| $ \begin{array}{cccccccccccccccccccccccccccccccccccc$ | .67 | .00 |
| $ \begin{array}{cccccccccccccccccccccccccccccccccccc$ | .30 | 15 |
| $ \begin{array}{cccccccccccccccccccccccccccccccccccc$ | .53 | 04 |
| $ \begin{array}{cccccccccccccccccccccccccccccccccccc$ | 95 | .09 |
| | 98 | .09 |
| $ \begin{array}{cccccccccccccccccccccccccccccccccccc$ | .08 | 05 |
| $ \begin{array}{cccccccccccccccccccccccccccccccccccc$ | .27 | .10 |
| $ \begin{array}{cccccccccccccccccccccccccccccccccccc$ | 08 | 03 |
| $ \begin{array}{cccccccccccccccccccccccccccccccccccc$ | .08 | 14 |
| $ \begin{array}{cccccccccccccccccccccccccccccccccccc$ | 14 | 11 |
| $ \begin{array}{cccccccccccccccccccccccccccccccccccc$ | .25 | 03 |
| $ \begin{array}{cccccccccccccccccccccccccccccccccccc$ | .22 | 02 |
| $ \begin{array}{cccccccccccccccccccccccccccccccccccc$ | .14 | 14 |
| $\begin{array}{cccccccccccccccccccccccccccccccccccc$ | .13 | 09 |
| $ \begin{array}{cccccccccccccccccccccccccccccccccccc$ | .29 | .04 |
| $ \begin{array}{cccccccccccccccccccccccccccccccccccc$ | .25 | 06 |
| $ \begin{array}{cccccccccccccccccccccccccccccccccccc$ | .02 | 11 |
| $ \begin{array}{cccccccccccccccccccccccccccccccccccc$ | .08 | 03 |
| $ \begin{array}{cccccccccccccccccccccccccccccccccccc$ | 32 | .36 |
| 8 8 1 1768.80173 .11 7 .94 .25 3502.60798 .14 5 8 8 0 1768.80173 .15 4 .94 .25 3502.60798 .28 * 9 0 9 916.25780 .03 23 .11 .01 2501.95934 .07 6 9 1 9 916.29353 .06 22 .06 07 2502.04076 .09 6 9 2 8 1074.76294 .03 33 06 14 2677.28337 .07 8 9 2 8 1075.90944 .06 26 .07 08 2679.53346 .05 6 9 3 7 1211.18563 .06 22 .22 13 2894.70558 .02 6 9 3 6 1279.79751 .03 21 49 .11 3122.96940 .07 5 9 4 5 1355.19908 .05 23 .05 08 | 69 | 27 |
| 880 $1/68.80173$.154.94.25 3502.60798 .28*909916.25780.0323.11.01 2501.95934 .076919916.29353.0622.06 07 2502.04076 .0969181074.76294.03.33 06 14 2677.28337 .0789281075.90944.06.26.07 08 2679.53346 .0569271198.19950.03.31 30 16 2808.28013 .0369361279.79751.03.21 49 16 2894.70558 .0269461334.47936.06.20 02 .01 2969.72197 .0579451355.19908.05.23.05 08 2986.25894 .0379551466.01848.0813.17.113122.96940.0759541468.61209.05.25.41 12 3124.71634.0569631619.05608.05.20.56.213300.88270.0979731794.37509.146.15.343502.42622.1229721794.38018.0710 $-$ | 66 | .17 |
| $\begin{array}{cccccccccccccccccccccccccccccccccccc$ | 66 | .17 |
| 919916.29353.06 22 06 07 2502.04076 0969181074.76294.0333 06 14 2677.28337 .0789281075.90944.0626.07 08 2679.53346 .0569271198.19950.0331 30 16 2808.28013 .0369361279.79751.0321 49 16 2894.70558 .0269461334.47936.0620 02 .01 2969.72197 .0579451355.19908.0523.05 08 2986.25894 .0379551466.01848.0813.17.11 3122.96940 .0759541468.61209.0525.41 12 3124.71634 .0569631619.05608.0520.56.21 3300.88270 .0979721794.38018.0710 31 27 3502.42622 .1249821989.35209.1001 68 .14 3723.80350 .30*9912200.40578.30*.52.43 3960.90911 .46*9902200.40578.15 | .18 | 02 |
| 918 $10/4, 76294$.0333 06 14 2677.28537 $.07$ 8928 1075.90944 .0626.07 08 2679.53346 .056927 1198.19950 .0331 30 16 2808.28013 .036936 1279.79751 .03 21 49 16 2894.70558 .026946 1334.47936 .06 20 02 .01 2969.72197 .057945 1355.19908 .05 23 .05 08 2986.25894 .037945 1355.19908 .05 23 .05 08 2986.25894 .037955 1466.01848 .08 13 .17.11 3122.96940 .075954 1468.61209 .05 25 .41 12 3124.71634 .056963 1619.05608 .05 20 .56.21 3300.88270 .097973 1794.37509 .146.15.34 3502.42260 .122972 1794.38018 .0710 31 27 3502.42622 .124982 1989.35209 1.00 1 68 .14 3723.80350 .30*99 <td< td=""><td>.04</td><td>07</td></td<> | .04 | 07 |
| $\begin{array}{cccccccccccccccccccccccccccccccccccc$ | 16 | 04 |
| $\begin{array}{cccccccccccccccccccccccccccccccccccc$ | 16 | 13 |
| 9571211.18365.062222 15 2829.36105 .0469361279.79751.0321 49 16 2894.70558 .0269461334.47936.0620 02 .01 2969.72197 .0579451355.19908.0523.05 08 2986.25894 .0379551466.01848.0813.17.113122.96940.0759541468.61209.0525.41 12 3124.71634.0569641618.89618.0814.02 16 3300.78714.0959631619.05608.0520.56.213300.88270.0979731794.37509.146.15.343502.42620.1229721794.38018.0710 31 27 3502.42622.1249821989.352091.001 68 .143723.80350.30*9902200.40578.30*.52.433960.90911.46*9902200.40578.153.52.433960.90911.234100101109.78693.0814.01 15 <td>22</td> <td>20</td> | 22 | 20 |
| 956 $12/9, 19/31$ $.03$ 21 49 16 $2894, 70338$ $.02$ 6946 1334.47936 $.06$ 20 02 $.01$ $2969,72197$ $.05$ 7945 1355.19908 $.05$ 23 $.05$ 08 2986.25894 $.03$ 7955 1466.01848 $.08$ 13 $.17$ $.11$ 3122.96940 $.07$ 5954 1468.61209 $.05$ 25 $.41$ 12 3124.71634 $.05$ 6964 1618.89618 $.08$ 14 $.02$ 16 3300.78714 $.09$ 5963 1619.05608 $.05$ 20 $.56$ $.21$ 3300.88270 $.09$ 7973 1794.37509 $.14$ 6 $.15$ $.34$ 3502.42260 $.12$ 2972 1794.38018 $.07$ 10 31 27 3502.42622 $.12$ 4982 1989.35209 1.00 1 68 $.14$ 3723.80350 $.30$ *990 2200.40578 $.30$ * $.52$ $.43$ 3960.90911 $.46$ *990 2200.40578 $.15$ 3 $.52$ $.43$ 3960.90911 $.23$ 410010 1109.78693 $.08$ 14 $.01$ 15 | .45 | 05 |
| $\begin{array}{cccccccccccccccccccccccccccccccccccc$ | .41 | 08 |
| 9431333.19908 $.03$ 23 $.03$ 06 2980.23894 $.03$ 7 9551466.01848 $.08$ 13 $.17$ $.11$ 3122.96940 $.07$ 5 9541468.61209 $.05$ 25 $.41$ 12 3124.71634 $.05$ 6 9641618.89618 $.08$ 14 $.02$ 16 3300.78714 $.09$ 5 9631619.05608 $.05$ 20 $.56$ $.21$ 3300.88270 $.09$ 7 9731794.37509 $.14$ 6 $.15$ $.34$ 3502.42260 $.12$ 2 9721794.38018 $.07$ 10 31 27 3502.42622 $.12$ 4 9821989.35209 1.00 1 68 $.14$ 3723.80350 $.30$ $*$ 9811989.35209 $.11$ 9 79 $.14$ 3723.80350 $.15$ 5 991 2200.40578 $.30$ $*$ $.52$ $.43$ 3960.90911 $.46$ $*$ 990 2200.40578 $.15$ 3 $.52$ $.43$ 3960.90911 $.23$ 4 10010 1109.78693 $.08$ 14 $.01$ 15 2693.95588 $.09$ 3 10110 1109.80210 $.06$ 19 $.21$ 01 | .29 | 08 |
| $\begin{array}{cccccccccccccccccccccccccccccccccccc$ | _ 33 | - 12 |
| $\begin{array}{cccccccccccccccccccccccccccccccccccc$ | 33 | 12 |
| $\begin{array}{cccccccccccccccccccccccccccccccccccc$ | .05 | 13 |
| $\begin{array}{cccccccccccccccccccccccccccccccccccc$ | 47 | 01 |
| $\begin{array}{cccccccccccccccccccccccccccccccccccc$ | 02 | 54 |
| 9 8 2 1989.35209 1.00 1 68 .14 3723.80350 .30 * 9 8 1 1989.35209 1.10 1 68 .14 3723.80350 .30 * 9 8 1 1989.35209 .11 9 79 .14 3723.80350 .15 5 9 9 1 2200.40578 .30 * .52 .43 3960.90911 .46 * 9 9 0 2200.40578 .15 3 .52 .43 3960.90911 .23 4 10 0 10 1109.78693 .08 14 .01 15 2693.95588 .09 3 10 1 10 1109.80210 .06 19 .21 01 2693.99334 .11 6 10 1 9 1287.73457 .08 18 32 28 2891.35965 .09 4 | -1.54 | 25 |
| 9 8 1 1989.35209 .11 9 79 .14 3723.80350 .15 5 9 9 1 2200.40578 .30 * .52 .43 3960.90911 .46 * 9 9 0 2200.40578 .15 3 .52 .43 3960.90911 .46 * 9 9 0 2200.40578 .15 3 .52 .43 3960.90911 .23 4 10 0 10 1109.78693 .08 14 .01 15 2693.95588 .09 3 10 1 10 1109.80210 .06 19 .21 01 2693.99334 .11 6 10 1 9 1287.73457 .08 18 32 28 2891.35965 .09 4 | _ 24 | _ 54 |
| 9 9 1 2200.40578 .30 * .52 .43 3960.90911 .46 * 9 9 0 2200.40578 .15 3 .52 .43 3960.90911 .23 4 10 0 10 1109.78693 .08 14 .01 15 2693.95588 .09 3 10 1 10 1109.80210 .06 19 .21 01 2693.99334 .11 6 10 1 9 1287.73457 .08 18 32 28 2891.35965 .09 4 | - 30 | .54 - 54 |
| 9 9 0 2200.40578 .15 3 .52 .43 3960.90911 .23 4 10 0 10 1109.78693 .08 14 .01 15 2693.95588 .09 3 10 1 10 1109.80210 .06 19 .21 01 2693.99334 .11 6 10 1 9 1287.73457 .08 18 32 28 2891.35965 .09 4 | 1.73 | .54 |
| $ \begin{array}{cccccccccccccccccccccccccccccccccccc$ | 1.73 | 11 |
| $ \begin{array}{cccccccccccccccccccccccccccccccccccc$ | - 21 | - 29 |
| 10 1 9 1287.73457 .08 18 32 28 2891.35965 .09 4 | .03 | 04 |
| | 37 | 12 |
| 10 2 9 1288.26732 .06 27 .14 03 2892.49206 .05 6 | 20 | 10 |
| 10 2 8 1433.02898 .11 130601 3047.09222 .08 5 | 27 | 16 |
| 10 3 8 1440.28795 .06 24 .2611 3060.10388 .05 7 | .46 | 10 |
| 10 3 7 1534,36830 .10 107007 3151,80967 .11 5 | .56 | 06 |
| 10 4 7 1574.44940 .06 213217 3210.49921 .05 7 | .47 | 10 |
| 10 4 6 1611.65347 .11 120710 3241.48621 .09 5 | .52 | 02 |
| 10 5 6 1709.54016 .07 1619 .40 3366.51770 .06 8 | 85 | 16 |
| 10 5 5 1716.19997 .10 10 .89 .45 3371.09888 .09 5 | .12 | 10 |
| 10 6 5 1862.45478 .08 130110 3544.33419 .12 6 | 66 | 28 |
| 10 6 4 1863.01995 .12 10 .38 .15 3544.67626 .14 3 | .07 | .06 |
| 10 7 4 2037.93390 .09 8 .3505 3746.24525 .13 5 | 91 | 12 |
| 10 7 3 2037.96234 .15 8 .83 .34 3746.26090 .16 4 | 64 | 08 |
| 10 8 3 2233.60249 .13 6 .0416 3968.60891 .22 3 | 27 | 54 |
| 10 8 2 2233.60330 .14 3 .00 .15 3968.61053 1.00 1 | .91 | .28 |
| 10 9 2 2445.95489 .14 4 -1.28 .39 4207.29643 .38 2 | 1.88 | -1223.86 |
| 10 9 1 2445.95489 .28 * -1.29 .39 4207.29643 .76 * | 1.87 | -1223.86 |

Table 3 (continued)

| J | Ka | K _c | (000) State | | | (010) State | | | | | | |
|----|--------|----------------|------------------|-------------------|------------------|--------------|------------|------------------|-------------------|-------------|-------------------------|------------|
| | | | $E_{\rm obs}/hc$ | $\delta_{ m obs}$ | N_{t} | Obs – Cal | Obs – [11] | $E_{\rm obs}/hc$ | $\delta_{ m obs}$ | $N_{\rm t}$ | Obs – Cal | Obs – [11] |
| 1 | | | 2 | 3 | 4 | 5 | 6 | 2 | 3 | 4 | 5 | 6 |
| 10 | 10 | 1 | 2671.63141 | .32 | 3 | 93 | 1.56 | 4458.50430 | .42 | 2 | -1.36 | -9321.65 |
| 10 | 10 | 0 | 2671.63141 | .64 | * | 93 | 1.56 | 4458.50430 | .82 | * | -1.36 | -9321.65 |
| 11 | 0 | 11 | 1321.45464 | .10 | 14 | .61 | .30 | 2903.84381 | .10 | 3 | .24 | .01 |
| 11 | 1 | 11 | 1321.45993 | .13 | 9 | 45 | 59 | 2903.86072 | .12 | 3 | 17 | 20 |
| 11 | 1 | 10 | 1518.54253 | .09 | 16 | 07 | 05 | 3122.94152 | .09 | 4 | .06 | .13 |
| 11 | 2 | 10 | 1518.78765 | .13 | 10 | 33 | 34 | 3123.50588 | .12 | 3 | 48 | 13 |
| 11 | 2 | 9 | 1684.43939 | .09 | 14 | .13 | 05 | 3302.22934 | .09 | 5 | 09 | 08 |
| 11 | 3 | 9 | 1688.28634 | .12 | 10 | .42 | 10 | 3309.74530 | .11 | 4 | .54 | 05 |
| 11 | 5 | 0 | 1825 48675 | .09 | 13 | 08 | 04 | 3429.72308 | .08 | 3 | .20 | 09 |
| 11 | 4 | 8 7 | 1894 19573 | .14 | 14 | =.31 - 40 | 03 | 3472.73127 | .13 | 4 | .14 | 20 |
| 11 | + 5 | 7 | 1076 20750 | .10 | 8 | 40 | 04 | 3633 45135 | .08 | 4 | -2.19 | 05 |
| 11 | 5 | 6 | 1990 85695 | 10 | 12 | 95 | .27 | 3643 79517 | 14 | - - | 08 | - 15 |
| 11 | 6 | 6 | 2129 89790 | 11 | 8 | - 60 | .15 | 3811 64173 | 17 | 3 | - 04 | .15 |
| 11 | 6 | 5 | 2131 56623 | 11 | 8 | 51 | 03 | 3812.65798 | .17 | 2 | -1 74 | -22.02 |
| 11 | 7 | 5 | 2305.19472 | .13 | 4 | -1.03 | -1.28 | 4013.61029 | .27 | 2 | .73 | 32 |
| 11 | 7 | 4 | 2305.30620 | .15 | 7 | .17 | .00 | 4013.67122 | .22 | 3 | .87 | 21 |
| 11 | 8 | 4 | 2501.3418 e | .24 | 2 | 4.27 | -110.93 | 4236.76638 | .22 | 2 | .24 | 37 |
| 11 | 8 | 3 | 2501.34271 | .35 | 2 | .41 | -114.61 | 4236.76869 | .25 | 3 | .17 | 26 |
| 11 | 9 | 3 | 2714.86265 | .58 | * | .94 | -1.15 | 4476.93274 | .86 | * | 1.01 | -937.86 |
| 11 | 9 | 2 | 2714.86265 | .29 | 2 | .81 | -2.23 | 4476.93274 | .43 | 2 | .95 | -937.86 |
| 11 | 10 | 2 | 2942.31325 | .62 | * | 35 | | 4730.22092 | 1.10 | * | 94 | |
| 11 | 10 | 1 | 2942.31325 | .31 | 2 | 35 | | 4730.22092 | .55 | 2 | 94 | |
| 11 | 11 | 1 | 3180.40686 | .90 | * | 29 | | 4993.04184 | 2.00 | * | .43 | |
| 11 | 11 | 0 | 3180.40686 | .45 | 2 | 29 | | 4993.04184 | 1.00 | 1 | .43 | |
| 12 | 0 | 12 | 1551.20192 | .18 | 9 | .16 | 20 | 3131.57453 | .15 | 3 | 56 | 65 |
| 12 | 1 | 12 | 1551.20495 | .14 | 12 | .45 | 01 | 3131.58354 | .12 | 3 | .14 | .15 |
| 12 | 1 | 11 | 1767.22493 | .17 | 7 | 56 | 18 | 3372.13535 | .14 | 3 | .03 | 17 |
| 12 | 2 | 11 | 1767.33885 | .13 | 11 | .03 | .20 | 3372.41748 | .12 | 4 | 08 | .07 |
| 12 | 2 | 10 | 1952.67784 | .16 | 8 | .12 | 31 | 35/3.6/535 | .18 | 2 | .02 | 09 |
| 12 | 3 | 10 | 1954.65355 | .11 | 12 | .46 | 01 | 35//.86092 | .11 | 4 | .18 | .02 |
| 12 | 3 | 9 | 2099.30272 | .24 | 0 | 42 | .17 | 3723.90382 | 1.00 | 1 | 59 | 08 |
| 12 | 4 | 9 | 2110.30478 | .13 | 10 | 21 | 08 | 3/33.43/91 | .10 | 4 | 10 | .01 |
| 12 | 4 | 8 | 2200.43789 | .22 | 3 4 | 38 | 99 | 3923 0464 e | 17 | 4 | _4 91 | 09 |
| 12 | 5 | 7 | 2203.43902 | 24 | 3 | .03 | - 62 | 3943 49842 | 1.00 | 1 | -1.04 | - 08 |
| 12 | 6 | 7 | 2420 88534 | 12 | 7 | - 64 | 14 | 4102 41570 | 32 | 2 | 2.39 | 372.60 |
| 12 | 6 | 6 | 2425.12367 | .16 | 4 | .76 | 03 | | | _ | , | |
| 12 | 7 | 6 | 2595.97487 | .16 | 3 | -1.05 | 32 | 4304.29645 | 1.55 | 1 | 47 | 643.85 |
| 12 | 7 | 5 | 2596.33807 | .19 | 3 | .25 | 23 | 4304.49920 | 2.00 | 1 | 1.93 | 632.98 |
| 12 | 8 | 5 | 2792.33816 | .29 | 2 | .34 | -2.84 | 4528.01177 | 1.64 | 2 | 2.41 | 744.87 |
| 12 | 8 | 4 | 2792.35668 | .20 | 3 | -1.37 | .45 | 4528.01615 | 2.00 | 1 | -3.47 | |
| 12 | 9 | 4 | 3006.85190 | .58 | 2 | 1.14 | 4.70 | 4769.50062 | .49 | 2 | .24 | |
| 12 | 9 | 3 | 3006.8568 e | 1.00 | 1 | 5.23 | 8.76 | 4769.50062 | .98 | * | 13 | |
| 12 | 10 | 3 | 3235.96129 | .40 | 2 | 28 | | 5024.77323 | .61 | 2 | -1.85 | |
| 12 | 10 | 2 | 3235.96129 | .80 | * | 30 | | 5024.77323 | 1.22 | * | -1.86 | |
| 12 | 11 | 2 | 3476.34286 | .45 | 2 | 1.09 | | 5290.20846 | .72 | 2 | .14 | |
| 12 | 11 | 1 | 3476.34286 | .90 | * | 1.09 | | 5290.20846 | 1.44 | * | .14 | |
| 12 | 12 | 1 | 3724.76172 | .64 | 2 | 69 | | | | | | |
| 12 | 12 | 12 | 3/24./61/2 | .64 | * 10 | 69 | 10 | 2277 00(00 | 21 | 2 | 07 | 16 |
| 13 | 0 | 13 | 1/98.96531 | .18 | 10 | .30 | 18 | 33//.09600 | .21 | 2 | .8/ | .46 |
| 13 | 1 | 13 | 1/20.20208 | .23 | o o | -1.14 | -1.29 | 3638 07564 | .22 | 2 | 13 | 09 |
| 13 | 1 2 | 12 | 2033.70007 | .22 | 0 5 | .13 | .3/ | 3630 11605 | .10 | 2 | .00 | 13 |
| 13 | 2 | 12 | 2033.01093 | . <i>∠1</i> 15 | 5 10 | -1.03 | -1.29 | 3861 71640 | .20 | 2 | 24 84 | 13 |
| 13 | 2 | 11 | 2230.03200 | .15 | 5 | 22 | _ 04 | 3863 00080 | 10 | 2 2 | .0 4 - 81 | .04 |
| 13 | 3 | 10 | 2406 77011 | 15 | 7 | 1 36 | 22 | 4038 65464 | 15 | 2 | 14 | - 12 |
| 13 | 4 | 10 | 2416 78334 | 20 | 4 | .00 | .22 | 4057 66668 | 1.00 | 1 | 46 | 13 |
| 13 | 4 | 9 | 2527.68873 | .20 | 3 | 44 | 14 | | 1.50 | | | .15 |
| 13 | 5 | 9 | 2575.97153 | .19 | 3 | -1.08 | 22 | | | | | |
| 13 | 5 | 8 | 2622.00999 | .16 | 2 | .84 | .72 | | | | | |

| J | Ka | $K_{\rm c}$ | (000) State | | | | | (010) State | | | | |
|----------|--------|-------------|------------------|-------------------|-------------|-----------|------------|------------------|-------------------|-------------|------------|----------|
| | | | $E_{\rm obs}/hc$ | $\delta_{ m obs}$ | $N_{\rm t}$ | Obs – Cal | Obs – [11] | $E_{\rm obs}/hc$ | $\delta_{ m obs}$ | $N_{\rm t}$ | Obs – Cal | Obs-[11] |
| 1 | | | 2 | 3 | 4 | 5 | 6 | 2 | 3 | 4 | 5 | 6 |
| 13 | 6 | 8 | 2734.89689 | .30 | 3 | -3.11 | .92 | | | | | |
| 13 | 6 | 7 | 2744.34538 | .19 | 3 | .33 | 18 | 4422.23241 | .38 | 2 | 1.17 | 529.21 |
| 13 | 7 | 7 | 2910.03916 | .27 | 3 | .32 | -1.42 | | | | | |
| 13 | 7 | 6 | 2911.06386 | .21 | 3 | .03 | .40 | 4618.63276 | 1.44 | 2 | 26 | |
| 13 | 8 | 6 | 3106.37348 | .42 | 2 | .16 | -14.52 | | | | | |
| 13 | 8 | 5 | 3106.44666 | .24 | 3 | 1.02 | 1.39 | | | | | |
| 13 | 9 | 5 | 3321.64477 | 1.00 | 1 | -1.42 | | 5004 (0270 | 1.70 | • | 1.07 | |
| 13 | 9 | 4 | 3321.64/60 | .39 | 2 | -2.17 | | 5084.68270 | 1.73 | 2 | 1.07 | |
| 13 | 10 | 4 | 3552.25/9/ | 1.54 | | 1.27 | | 5341.8084 e | 1.38 | ۳ | 8.10 | |
| 13 | 10 | 3 | 3794 80192 | 1.06 | * | 1.14 | | 5609 73587 | 2.09 | * | 8.04 88 | |
| 13 | 11 | 2 | 3794 80192 | 53 | 2 | .24 | | 5609.73587 | 2.00 | 1 | .00 | |
| 13 | 12 | 2 | 4046 06474 | 3 16 | * | 22 | | 5007.15501 | 1.00 | 1 | .00 | |
| 13 | 12 | 1 | 4046.06474 | 1.58 | 2 | .22 | | | | | | |
| 13 | 13 | 1 | 4302.84613 | 1.64 | * | .41 | | | | | | |
| 13 | 13 | 0 | 4302.84613 | .82 | 2 | .41 | | | | | | |
| 14 | 0 | 14 | 2064.67217 | .26 | 6 | 07 | 08 | 3640.34209 | 1.00 | 1 | .42 | 18 |
| 14 | 1 | 14 | 2064.67325 | .23 | 8 | .47 | .34 | 3640.34329 | .27 | 3 | 64 | 15 |
| 14 | 1 | 13 | 2318.11350 | .31 | 5 | 60 | 17 | 3923.44772 | .33 | 1 | 64 | -1.30 |
| 14 | 2 | 13 | 2318.13791 | .52 | 3 | -1.24 | 74 | 3923.52235 | 1.00 | 1 | .89 | .40 |
| 14 | 2 | 12 | 2540.66999 | 1.00 | 1 | 41 | -1.16 | | | | | |
| 14 | 3 | 12 | 2541.17273 | .23 | 4 | 13 | .08 | 4167.85024 | 1.00 | 1 | -1.13 | 02 |
| 14 | 3 | 11 | 2729.76584 | .31 | 2 | 15 | -2.96 | | | | | |
| 14 | 4 | 11 | 2735.41974 | .27 | 2 | 1.70 | .53 | 4378.55945 | 1.00 | 1 | 1.27 | .16 |
| 14 | 4 | 10 | 28/3.3231/ | 1.00 | 1 | 2.55 | .10 | | | | | |
| 14 | 5 | 10 | 2906.86346 | .37 | 2 | 1.23 | 14.46 | | | | | |
| 14 | 5 | 9 | 29/0.08/95 | .39 | 2 | 1.60 | 18.90 | | | | | |
| 14 | 6 | 9 | 3089 89726 | .41 | 2 | .90 | 18.90 | | | | | |
| 14 | 7 | 8 | 3247 06781 | 36 | 3 | - 55 | 20.77 | 4954 60453 | 56 | 2 | -1 54 | |
| 14 | 7 | 7 | 3249 63173 | 44 | 2 | -1.52 | 20.77 | 1991.00199 | .50 | 2 | 1.51 | |
| 14 | 8 | 7 | 3443.20007 | .47 | 2 | 14 | | 5178.66720 | 1.53 | 2 | -1.50 | |
| 14 | 8 | 6 | 3443.42357 | .50 | 2 | .34 | | | | | | |
| 14 | 9 | 6 | 3658.96746 | .46 | 2 | .97 | | | | | | |
| 14 | 9 | 5 | 3658.97888 | .62 | 2 | -1.39 | | | | | | |
| 14 | 10 | 5 | 3890.87507 | .53 | 2 | 43 | | 5680.9221 e | 1.79 | 1 | -12.57 | |
| 14 | 10 | 4 | 3890.87507 | 1.06 | * | -1.05 | | 5680.9221 e | 3.58 | * | -12.84 | |
| 14 | 11 | 4 | 4135.42990 | .94 | 2 | 1.15 | | 5951.22574 | 1.00 | 1 | 17 | |
| 14 | 11 | 3 | 4135.42990 | 1.88 | * | 1.13 | | 5951.22574 | 2.00 | * | 18 | |
| 14 | 12 | 3 | 4389.38810 | .68 | 2 | .06 | | | | | | |
| 14 | 12 | 2 | 4389.38810 | 1.30 | т Э | .06 | | | | | | |
| 14 | 13 | 1 | 4649.62016 | 3.34 | * | 1.14 | | | | | | |
| 14 | 13 | 1 | 4912 92118 | 1.00 | 2 | 1.14 | | | | | | |
| 14 | 14 | 0 | 4912.92118 | 2.00 | * | 1.08 | | | | | | |
| 15 | 0 | 15 | 2348.24641 | .27 | 7 | .31 | 07 | 3921.24784 | .28 | 2 | .40 | 06 |
| 15 | 1 | 15 | 2348.2456 e | .33 | 5 | 81 | 95 | 3921.24784 | .56 | * | 96 | 26 |
| 15 | 1 | 14 | 2620.20327 | .49 | 2 | 1.98 | 30.27 | 4225.51938 | 1.00 | 1 | .64 | 26 |
| 15 | 2 | 14 | 2620.21433 | 1.00 | 1 | .79 | 26.33 | 4225.55689 | 1.00 | 1 | 51 | 18 |
| 15 | 2 | 13 | 2860.65722 | .27 | 2 | .67 | .51 | 4488.52335 | 1.00 | 1 | 15 | .03 |
| 15 | 3 | 13 | 2860.90859 | 1.00 | 1 | -1.93 | .09 | | | | | |
| 15 | 3 | 12 | 3068.81808 | 1.45 | 1 | 2.56 | -4.10 | | | | | |
| 15 | 4 | 12 | 3071.92687 | 1.00 | 1 | .96 | .00 | | | | | |
| 15 | 4 | 11 | 3235.24224 | 1.00 | 1 | -3.83 | -1.42 | | | | | |
| 15 | 5 | 10 | 3352.80902 | 1.32 | 2 | -1.17 | -20.19 | | | | | |
| 15 | 6 | 10 | 3429.03500 | .08 | 3 | 09 | 35.20 | | | | | |
| 15 15 | 6 7 | 9 | 3401.9001/ | .37 | 2 1 | 1.49 | | | | | | |
| 15 | 7 | 7 8 | 3612 30004 | 62 | 1 2 | .34 | | | | | | |
| 15 | , 8 | 8 | 3802 54963 | .02 60 | 2 | - 74 | | | | | | |
| 15 | 8 | 7 | 3803.15922 | .53 | 2 | .51 | | 5537.83127 | 1.00 | 1 | .85 | |
| | | | | | | | | | | | | |

Table 3 (continued)

| J | Ka | Kc | (000) State | | | (010) State | | | | | | |
|----|----|----|------------------|-------------------|------------------|-------------|------------|------------------|--------------------|-------------|-----------|------------|
| | | | $E_{\rm obs}/hc$ | $\delta_{ m obs}$ | N_{t} | Obs – Cal | Obs – [11] | $E_{\rm obs}/hc$ | $\delta_{\rm obs}$ | $N_{\rm t}$ | Obs – Cal | Obs – [11] |
| 1 | | | 2 | 3 | 4 | 5 | 6 | 2 | 3 | 4 | 5 | 6 |
| 15 | 9 | 7 | 4018.52907 | .71 | 2 | 1.65 | | | | | | |
| 15 | 9 | 6 | 4018.57347 | .63 | 2 | .48 | | 5781.61591 | 1.60 | 1 | .30 | |
| 15 | 10 | 6 | 4251.49245 | .77 | 2 | 54 | | | | | | |
| 15 | 10 | 5 | 4251.49683 | .68 | 2 | 1.28 | | | | | | |
| 15 | 11 | 5 | 4497.86311 | 1.38 | * | 56 | | | | | | |
| 15 | 11 | 4 | 4497.86311 | .69 | 2 | 67 | | | | | | |
| 15 | 12 | 4 | 4754.34659 | 2.24 | * | 03 | | | | | | |
| 15 | 12 | 3 | 4/54.34659 | 1.12 | 2 | 04 | | | | | | |
| 15 | 13 | 3 | 5017.83693 | 1.00 | τ · | /3 | | | | | | |
| 15 | 13 | 2 | 5285 25001 | .03 | * | 75 | | | | | | |
| 15 | 14 | 2 | 5285.25091 | 3 27 | 2 | -2.81 | | | | | | |
| 15 | 14 | 1 | 5553 35746 | 2 34 | * | -2.81 | | | | | | |
| 15 | 15 | 0 | 5553 35746 | 1 17 | 2 | 88 | | | | | | |
| 16 | 0 | 16 | 2649 60314 | 41 | 2 | - 69 | -1.67 | 4219 74127 | 2.00 | * | 70 | - 04 |
| 16 | 1 | 16 | 2649.60430 | .30 | 4 | .30 | 51 | 4219.74127 | 1.00 | 1 | 27 | 04 |
| 16 | 2 | 14 | 3197.97471 | 1.00 | 1 | -1.72 | -1.29 | | | - | | |
| 16 | 3 | 14 | 3198.0793 e | .33 | 2 | -27.33 | 47 | | | | | |
| 16 | 4 | 13 | 3425.91891 | .53 | 2 | 2.32 | 30.45 | | | | | |
| 16 | 5 | 12 | 3625.88154 | 1.06 | 2 | -2.96 | 84.83 | | | | | |
| 16 | 6 | 11 | 3807.33419 | 1.44 | 3 | -3.09 | | | | | | |
| 16 | 7 | 10 | 3988.16976 | .72 | 2 | 1.70 | | | | | | |
| 16 | 7 | 9 | 3999.84787 | .81 | 2 | -2.06 | | | | | | |
| 16 | 8 | 9 | 4184.11632 | .80 | 2 | 1.78 | | | | | | |
| 16 | 9 | 8 | 4400.04331 | .90 | 2 | 1.14 | | | | | | |
| 16 | 9 | 7 | 4400.17492 | .76 | 2 | 50 | | | | | | |
| 16 | 10 | 7 | 4633.78814 | 2.72 | 2 | 1.88 | | | | | | |
| 16 | 10 | 6 | 4633.79434 | .90 | 2 | 90 | | | | | | |
| 16 | 11 | 6 | 4881.75047 | .87 | 2 | 1.54 | | | | | | |
| 16 | 11 | 5 | 4881./4919 | 1.50 | 1 | 21 | | | | | | |
| 16 | 12 | 3 | 5140.55300 | .01 | * | -3.09 | | | | | | |
| 16 | 12 | 4 | 5407 09812 | 1.02 | 2 | -1.52 | | | | | | |
| 16 | 13 | 3 | 5407.09812 | 2.58 | * | -1.52 | | | | | | |
| 16 | 14 | 3 | 5678.37599 | 1.00 | 2 | 1.38 | | | | | | |
| 16 | 14 | 2 | 5678.37599 | 2.00 | * | 1.38 | | | | | | |
| 16 | 15 | 2 | 5951.3247 e | 3.35 | 2 | 4.84 | | | | | | |
| 16 | 15 | 1 | 5951.3247 e | 6.70 | * | 4.84 | | | | | | |
| 16 | 16 | 1 | 6222.63276 | 1.35 | 2 | .34 | | | | | | |
| 16 | 16 | 0 | 6222.63276 | 2.70 | * | .34 | | | | | | |
| 17 | 0 | 17 | | | | | | 4535.74522 | 1.00 | 1 | .36 | 49 |
| 17 | 1 | 17 | | | | | | 4535.74522 | 2.00 | * | 45 | 49 |
| 17 | 4 | 13 | 4004.33559 | 1.00 | 1 | 54 | | | | | | |
| 17 | 5 | 12 | 4163.74021 | 1.00 | 1 | 1.05 | | | | | | |
| 17 | 6 | 11 | 4281.56959 | 1.66 | 2 | 41 | | | | | | |
| 17 | / | 10 | 4412.50661 | 1.54 | 2 | .10 | | | | | | |
| 17 | ð | 10 | 4307.3123 e | 1.00 | 1 | -7.44 | | | | | | |
| 17 | 0 | 8 | 4390.8707 6 | 1.00 | 1 | -1.65 | | | | | | |
| 17 | 10 | 8 | 5037 43758 | 94 | 2 | - 05 | | | | | | |
| 17 | 10 | 7 | 5037.46578 | 1.05 | 2 | .31 | | | | | | |
| 17 | 11 | 7 | 5286.73127 | 1.05 | 1 | -1.28 | | | | | | |
| 17 | 11 | 6 | 5286.73577 | 2.81 | 2 | 1.48 | | | | | | |
| 17 | 12 | 6 | 5547.63766 | 2.06 | * | .24 | | | | | | |
| 17 | 12 | 5 | 5547.63766 | 1.03 | 2 | .15 | | | | | | |
| 17 | 13 | 5 | 5817.00823 | 5.08 | * | .02 | | | | | | |
| 17 | 13 | 4 | 5817.00823 | 2.54 | 2 | .02 | | | | | | |
| 17 | 14 | 4 | 6091.88529 | 2.92 | * | 3.53 | | | | | | |
| 17 | 14 | 3 | 6091.88529 | 1.46 | 2 | 3.53 | | | | | | |
| 17 | 15 | 3 | 6369.3087 e | 2.32 | * | -12.49 | | | | | | |
| 17 | 15 | 2 | 6369.3087 e | 1.16 | 1 | -12.49 | | | | | | |

Table 3 (continued)

| J | $K_{\rm a}$ | $K_{\rm c}$ | (000) State | | | | | (010) State | | | | | | |
|----|-------------|-------------|------------------|-------------------|----|-----------|------------|------------------|-------------------|-------------|-----------|------------|--|--|
| | | | $E_{\rm obs}/hc$ | $\delta_{ m obs}$ | Nt | Obs – Cal | Obs – [11] | $E_{\rm obs}/hc$ | $\delta_{ m obs}$ | $N_{\rm t}$ | Obs – Cal | Obs - [11] | | |
| 1 | | | 2 | 3 | 4 | 5 | 6 | 2 | 3 | 4 | 5 | 6 | | |
| 17 | 16 | 2 | 6646.26112 | 6.82 | * | .28 | | | | | | | | |
| 17 | 16 | 1 | 6646.26112 | 3.41 | 1 | .28 | | | | | | | | |
| 17 | 17 | 1 | 6919.30188 | 3.02 | * | 15 | | | | | | | | |
| 17 | 17 | 0 | 6919.30188 | 1.51 | 1 | 15 | | | | | | | | |
| 18 | 7 | 12 | 4814.20190 | 2.02 | 1 | 2.11 | | | | | | | | |
| 18 | 8 | 11 | 5012.30542 | 1.62 | 1 | 47 | | | | | | | | |
| 18 | 9 | 10 | 5227.74643 | 1.04 | 1 | 52 | | | | | | | | |
| 18 | 11 | 8 | 5712.47021 | 1.19 | 1 | 55 | | | | | | | | |
| 18 | 11 | 7 | 5712.47521 | 1.09 | 1 | -1.21 | | | | | | | | |
| 18 | 12 | 7 | 5975.22046 | 2.87 | 1 | 1.13 | | | | | | | | |
| 18 | 12 | 6 | 5975.22046 | 5.74 | * | .80 | | | | | | | | |
| 18 | 13 | 6 | 6247.17369 | 2.00 | 1 | 07 | | | | | | | | |
| 18 | 13 | 5 | 6247.17369 | 4.00 | * | 08 | | | | | | | | |
| 18 | 14 | 5 | 6525.3893 e | 2.85 | 1 | 9.25 | | | | | | | | |
| 18 | 14 | 4 | 6525.3893 e | 5.70 | * | 9.25 | | | | | | | | |
| 18 | 15 | 4 | 6806.98239 | 1.73 | 1 | -1.66 | | | | | | | | |
| 18 | 15 | 3 | 6806.98239 | 3.46 | * | -1.66 | | | | | | | | |

Note. E_{obs}/hc , experimental term values in cm⁻¹; levels labeled with (e) in the next column are excluded from the fit (blended or very weak transitions: see the text). δ_{obs} , standard error of experimental value in 10^{-3} cm⁻¹. N_t , number of observed transitions used to define the energy level. Transitions to those levels indicated by * in this column were not resolved. The energy for such levels were fixed to those of observed coincident partners. Obs – Calc, difference between observed and calculated term values, in 10^{-3} cm⁻¹. Column 6: Obs – [11], difference between our observed term values and those of [11], in 10^{-3} cm⁻¹.

of $(0\ 0\ 0)$ and 11 in the case of $(0\ 1\ 0)$. In order to account for the anomalous centrifugal distortion at high-rotational energies, we use in the present study the generating function model [29,30] which was shown to allow more accurate calculations [28,31,36].

The reduced effective rotational Hamiltonian for an isolated vibration state of an asymmetric top molecule is written in the form

$$H_{\rm rot} = {}^{\rm d}H_{\rm rot} + {}^{\rm nd}H_{\rm rot}.$$
 (1)

The diagonal ${}^{d}H_{rot}$ and the non-diagonal ${}^{nd}H_{rot}$ parts of the Hamiltonian relative to $|J, K\rangle$ wavefunctions are expanded in the generating function **G** following [30]:

$${}^{\mathrm{d}}H_{v_1v_2v_3} = \sum_{nm} g_{nm} \mathbf{J}^{2n} \{ \mathbf{G}(\boldsymbol{\alpha}^{(J)}) \}^m, \tag{2}$$

$${}^{\mathrm{nd}}H_{v_1v_2v_3} = \sum_{nm} u_{nm} \mathbf{J}^{2n} \Big[\big(J_+^2 + J_-^2 \big), \Big\{ \mathbf{G}(\beta^{(J)}) \Big\}^m \Big]_+, \tag{3}$$

where the generating function is defined according to [29]

$$\mathbf{G} \equiv \mathbf{G}(\alpha^{(J)}) = (2/\alpha^{(J)}) \left\{ \sqrt{1 + \alpha^{(J)} J_z^2} - 1 \right\}.$$
 (4)

Relations of parameters with Watson constants and extrapolation properties of the model are discussed in [30]. The *J*-dependence of $\alpha^{(J)}$ in the generating function, is given by the development

$$\alpha^{(J)} = \sum_{n} \alpha_n \mathbf{J}^{2n}.$$
 (5)

Linear parameters g_{nm} , u_{nm} of the Hamiltonian expansion as well as non-linear parameters of the generating functions α_n and β_n for a given vibration state have been determined in a least squares fit to experimental rotational energies. Discrepancies between observed and calculated levels given in Table 4 (the column Obs-Cal in 10^{-3} cm⁻¹) and Figs. 1 and 2 (upper parts) show quite good agreement for the entire range of experimentally accessible rotational quantum numbers and also prove a consistency of our assignments. The root mean square (RMS) deviation between observed and calculated values is 0.0009 cm⁻¹ for 305 rotational levels of the (000) state with $J_{\text{max}} = 18$, $K_{\text{a max}} = 17$ and $0.0007 \,\text{cm}^{-1}$ for 198 rotational levels of the (010) state with $J_{\text{max}} = 17$, $K_{a max} = 11$. Several outliers labeled by (e) in Table 3 (7 observed levels for (000) state and 2 observed levels for the (010) state) corresponding to blended or very weak lines were excluded from the fit because the standard deviation δ_{obs} of their determination does not provide a reliable experimental error estimation in these cases.

Values of fitted Hamiltonian parameters for (000) and (010) states with their standard errors are given in Table 4. Note that it was possible to obtain a similar RMS deviation of a fit with considerably less number of parameters. However we prefer to obtain such a set of parameters which would provide an accurate description of experimental data, and simultaneously assure physically meaningful extrapolations consistent with so called

| able 4 | |
|---|--|
| itted values of the parameters of the effective Hamiltonian for (000) and (010) states of $H_2^{18}O$ | |

| Parameter | | (000) | | (010) | |
|------------------------|----------------------|---------------|-----------|-----------------|----------|
| | | Value | SE | Value | SE |
| E_{vv} | | | | 1588.27563 | |
| α ₀ | $\times 10^2$ | 4.477767422 | 0.031 | 1.415424762 | 0.0011 |
| α_1 | $\times 10^4$ | 1.067132336 | 0.0079 | -0.062152 | |
| α2 | $\times 10^{8}$ | 5.287635172 | 0.059 | 9.701555245 | 0.13 |
| α ₃ | $\times 10^{10}$ | 1.8235 | | -0.043206 | |
| α_4 | $\times 10^{15}$ | | | 2.0352 | |
| g_{10} | 2 | 11.87993852 | 0.0000057 | 11.88795872 | 0.000024 |
| g_{20} | $\times 10^{3}$ | -1.251811913 | 0.00014 | -1.399171908 | 0.00081 |
| g ₃₀ | $\times 10^{\prime}$ | 5.395683962 | 0.011 | 7.034297894 | 0.085 |
| g_{40} | $\times 10^{10}$ | -3.306976284 | 0.025 | -5.231815853 | 0.27 |
| g_{50} | $\times 10^{13}$ | 1.3445 | | 2.6523 | |
| g_{01} | | 15.651482 | 0.000039 | 18.8424396 | 0.00012 |
| g_{11} | ×10 ³ | 5.718235707 | 0.0017 | 7.699877632 | 0.0036 |
| <i>g</i> ₂₁ | ×10 ⁶ | -1.939333991 | 0.019 | 2.322002123 | 0.052 |
| g_{31} | $\times 10^9$ | 3.055332799 | 0.070 | 4.659577624 | 0.022 |
| g_{41} | $\times 10^{12}$ | 0.5186 | | -1.7178 | |
| g_{02} | $\times 10^{1}$ | 1.434570368 | 0.012 | 0.09866389008 | 0.0043 |
| g_{12} | $\times 10^4$ | 4.652862021 | 0.035 | -0.5687971522 | 0.0023 |
| g_{22} | $\times 10^{7}$ | 2.356069596 | 0.029 | 1.437138832 | 0.050 |
| g_{32} | $\times 10^{10}$ | 9.9666 | | -0.42314 | |
| g_{42} | $\times 10^{13}$ | -3.244 | | -0.44358 | |
| g_{03} | $\times 10^4$ | -5.838389273 | 0.053 | -0.05245 | |
| g_{13} | $\times 10^{\circ}$ | -1.763410418 | 0.021 | 1.058462072 | 0.0058 |
| g_{23} | $\times 10^{9}$ | -3.266178248 | 0.026 | 1.894 | |
| g_{33} | ×10 ¹² | -4.3966 | | | |
| g_{04} | $\times 10^{\circ}$ | 7958849337 | 0.0086 | -1.073564794 | 0.0027 |
| g_{14} | $\times 10^{9}$ | 1.01036748 | 0.067 | -0.009919217634 | 0.000036 |
| g_{24} | $\times 10^{12}$ | 8.1763 | | -8.592232485 | 0.12 |
| g_{05} | $\times 10^{8}$ | 1.085 | | 0.89053 | |
| g_{15} | $\times 10^{11}$ | 1.0223 | | 2.7816 | |
| g_{06} | $\times 10^{11}$ | -6.0545 | | -2.3 | |
| g_{07} | ×10 ¹⁴ | 5.328 | | | |
| β_0 | $\times 10^{2}$ | 4.396485501 | 0.11 | 3.749234274 | 0.057 |
| β_1 | $\times 10^4$ | -1.244393659 | 0.040 | | |
| β_2 | $\times 10^{7}$ | 3.185 | | | |
| u_{00} | | 1.320951929 | 0.0000056 | 1.398496406 | 0.000019 |
| u_{10} | $\times 10^4$ | -5.073963802 | 0.0019 | -5.789391333 | 0.0056 |
| u_{20} | $\times 10^{7}$ | 2.690239101 | 0.022 | 3.544197177 | 0.056 |
| u_{30} | $\times 10^{10}$ | -1.626609466 | 0.10 | -2.40024899 | 0.20 |
| u_{40} | $\times 10^{14}$ | 5.692388736 | 1.7 | 9.023953334 | 2.1 |
| u_{50} | $\times 10^{17}$ | | | 5.748 | |
| u_{01} | $\times 10^{3}$ | -1.297186101 | 0.0014 | -3.806116496 | 0.0050 |
| u_{11} | $\times 10^{6}$ | -1.000729809 | 0.016 | -2.324801482 | 0.051 |
| <i>u</i> ₂₁ | $\times 10^{9}$ | 0.521460084 | 0.052 | -1.759089563 | 0.17 |
| <i>u</i> ₃₁ | $\times 10^{12}$ | -0.047968 | | -4.9614 | |
| u_{02} | $\times 10^4$ | 0.2521821036 | 0.0014 | 1.000485109 | 0.0036 |
| <i>u</i> ₁₂ | $\times 10^{7}$ | 0.2299398745 | 0.0092 | 1.248833415 | 0.031 |
| <i>u</i> ₂₂ | $\times 10^{11}$ | -2.667826528 | 0.17 | 2.185222924 | 0.80 |
| <i>u</i> ₃₂ | $\times 10^{13}$ | | | 1.4641 | |
| u_{03} | ×10 ⁶ | -0.213229352 | 0.0032 | -1.636713199 | 0.0078 |
| <i>u</i> ₁₃ | $\times 10^9$ | 0.02828419899 | 0.0074 | -2.376925742 | 0.044 |
| <i>u</i> ₂₃ | $\times 10^{13}$ | 3.4056 | | | |
| u_{04} | $\times 10^8$ | 0.1297546064 | 0.0011 | 1.482109089 | 0.0081 |
| u_{14} | $\times 10^{12}$ | -1.0386 | | 1.8718 | |
| <i>u</i> ₀₅ | ×10 ¹¹ | -0.2062 | | -2.8332 | |

Note. All linear parameters g_{nm} , u_{nm} are in cm⁻¹. Non-linear parameters α_n and β_n are dimensionless.



Fig. 1. Comparison of observed levels included in the fit with calculations for the (000) state of H₂¹⁸O. The *upper part* corresponds to our calculations with the parameters of Table 4; the *lower part* corresponds to global PS predictions of [33]. The *abscissa* is defined as $J + (K_a/J)$ for J > 0 and zero for J = 0. Points at integer values in the horizontal axis represent *J*-behaviour of Obs – Calc for boundaries of a K_a manifold ($K_a = 0$ and $K_a = J$), whereas K_a -behaviour at a given *J* is represented by a series of points between two integer abscissa values.

"global calculations" (see the following sections for the related discussion) for a wider range of quantum numbers not yet accessible in current experimental spectra. To achieve this end we have introduced some higher order parameters whose values were constrained in a way to obtain a good agreement with global calculations outside the experimentally accessible range. These parameters correspond to the values without standard errors in Table 4.

4. Discussion: isotopic shifts and comparison with global calculations

It is instructive to compare our observations with the most accurate available global predictions from the molecular potential function by Partridge and Schwenke (PS) [33]. Discrepancies between our observed and calculated (PS) energy levels are plotted in Figs. 1b and 2b against rotational quantum numbers. The overall quali-



Fig. 2. Comparison of observed levels included in the fit with calculations for the $(0\,1\,0)$ state of $H_2^{18}O$. The *upper part* corresponds to our calculations with the parameters of Table 4; the *lower part* corresponds to global PS predictions of [33]. Notations are the same as in Fig. 1.

 Table 5

 Summary of comparisons of calculations with observed levels and with global predictions

| H ₂ ¹⁸ O | (000) State | (010) State | | | |
|---|-------------|-------------|--|--|--|
| (A) Fit of observed data using the G-function model (parameters of Table 4) | | | | | |
| J _{max} , K _{a max} | 18, 17 | 17, 11 | | | |
| N _{levels} (obs) | 305 | 198 | | | |
| RMS (Obs – Calc) $[cm^{-1}]$ | 0.0009 | 0.0007 | | | |
| (B) Comparison of our calculations with global predictions (PS) of [33] | | | | | |
| J _{max} , K _{a max} | 24, 24 | 22, 22 | | | |
| $E_{\rm rot\ max}/hc\ [{\rm cm}^{-1}]$ | 12460 | 12 793 | | | |
| N _{levels} (calc.) | 624 | 529 | | | |
| RMS (our-PS) $[cm^{-1}]$ | 0.26 | 0.21 | | | |



Fig. 3. Isotopic shifts [in cm⁻¹] for rotational term values versus K_a for fixed values of J for the (000) vibration state (*upper part*) and for the (010) vibration state (*lower part*). The vertical axis corresponds to $E_{J\cdot K_a} = E_{J\cdot K_a} (H_2^{16}O) - E_{J\cdot K_a} (H_2^{18}O)$. Dots correspond to series of J = 25, empty squares to J = 20, triangles to J = 15, and empty circles to J = 10.

tative agreement is surprisingly good. This gives another independent confirmation of a validity of our assignments for those levels where Table 3 shows large disagreements with previously determined energies. On the other hand this confirm remarkably good extrapolation properties of the potential function of [33]. It should be noticed that the global calculations of Partridge and Schwenke [33] are based on very careful ab initio calculations of the potential function with its subsequent optimization through a fit to spectroscopic data up to $J \le 5$. Thus for $J \ge 6$ the comparison relies on PS predictions. Note that for the same J values the agreement in K_a behaviour is considerably better for (0 0 0) state than for the (0 1 0) state. This suggests that a further optimization of a bending section of the molecular potential function using our new data should be still possible.

Global calculations based on a full rovibrational Hamiltonian like those performed by Partridge and

Schwenke [33] have the advantage of a "birds-eye" view of the entire set of rovibrational states. Though at present for medium values of quantum numbers they are not as accurate as best "local" calculations (based on an appropriately constructed effective Hamiltonian for an isolated vibration state or for a group of closely lying vibrational states), they can give more smooth and physically consistent extrapolations for high v, J, and K_a levels.

Comparisons given in Table 5 confirm the conclusion that results from our previous analysis of $H_2^{16}O$ spectra that the two types of calculations, "local" and "global" once, could be used as complementary tools in spectra analysis.

Our "local" model with parameters given in Table 4 allow to achieve rather accurate data reduction (in average better than 10^{-3} cm⁻¹) for presently available experimental data for (000) and (010) states as is seen in Figs. 1 and 2 (upper parts). On the other hand we can use global predictions to obtain parameter values giving smooth physically consistent extrapolations in a certain range of quantum numbers which is wider than that accessible experimentally. In our case the RMS deviation between our extrapolations and PS global predictions is only $0.26 \,\mathrm{cm}^{-1}$ for levels of the (000) state up to J, $K_{\mathrm{a}} = 24$ and 0.21 cm^{-1} for levels of the (010) state up to J, $K_a = 22$ which should correspond to a model accuracy at highrotational energies at $E_{\rm rot}/hc \sim 12,000 \,{\rm cm^{-1}}$. Note that such a good agreement would hardly be possible to achieve with a usual polynomial effective rotational Hamiltonian.

New experimental data and calculations allow us to study an isotopic shift of rotational levels as a function of $J, K_{\rm a}$ quantum numbers. The corresponding behaviour of shifts defined as $\Delta E_{J \cdot K_a} = E_{J \cdot K_a} (H_2^{16}O) - E_{J \cdot K_a} (H_2^{18}O)$ for fixed values J = 10, J = 15, J = 20, and J = 24 versus K_a in (000) and (010) vibration states is shown in Fig. 3. For those $H_2^{18}O$ energies which are not known experimentally the extrapolations with parameters of Table 4 are used to calculate isotopic shifts. Extensive tests of the model with available H2¹⁸O data and comparisons with the most accurate global calculations [33] suggest that the accuracy of extrapolations for the considered states should be of the order of $0.1-1 \,\mathrm{cm}^{-1}$ or better. Such errors are invisible on the scale of the Fig. 3. It is seen that for big K_a values the behaviour of isotopic shifts approaches a linear one in agreement with a proposal of [29]. The slopes of these asymptotics are nearly Jindependent in the considered range. This can be useful for the further assignment of high K_a transitions of rare isotopic species.

Acknowledgments

S.N.M. acknowledges support from the CNRS-RFBR PICS project No. 01-05-22002. S.N.M. and Vl.G.T. thank D. Schwenke for making available his global calculations and S. Tashkun for the collaboration in RITZ project and making available the corresponding code. The authors are grateful to M.Winnewisser, B. Winnewisser, and S. Klee for the help and valuable discussions. Vl.G.T. acknowledges support from the IDRIS computer centre of CNRS.

References

- P.E. Fraley, K. Narahari Rao, L.H. Jones, J. Mol. Spectrosc. 29 (1969) 312–317.
- [2] J.G. Williamson, K. Narahari Rao, L.H. Jones, J. Mol. Spectrosc. 40 (1971) 372–387.
- [3] J.W.C. Johns, A.R.W. McKeller, Can. J. Phys. 56 (1978) 737-743.
- [4] G. Guelachvili, J. Opt. Soc. Am. 73 (1983) 137-150.
- [5] C. Camy-Peyret, J.M. Flaud, N. Papineau, C.R. Acad. Sci. Paris B290 (1980) 537–540.
- [6] R.A. Toth, J. Opt. Soc. Am. B9 (1992) 462-482.
- [7] G. Steenbeckeliers, J. Bellet, C.R. Acad. Sci. Paris B273 (1971) 471–474.
- [8] F.C. De Lucia, P. Helminger, R.L. Cook, W. Gordy, Phys. Rev. A6 (1972) 1324–1326.
- [9] J.W.C. Johns, J. Opt. Soc. Am. B2 (1985) 1340-1354.
- [10] J. Kauppinen, E. Kyro, J. Mol. Spectrosc. 84 (1980) 405-423.
- [11] R.A. Toth, J. Mol. Spectrosc. 190 (1998) 379-396.
- [12] F. Matsushima, H. Nagase, T. Nakauchi, K. Takagi, J. Mol. Spectrosc. 193 (1999) 217–223.
- [13] R.A. Toth, J. Margolis, J. Mol. Spectrosc. 57 (1975) 236-245.
- [14] C. Camy-Peyret, J.M. Flaud, R.A. Toth, J. Mol. Spectrosc. 87 (1981) 233–241.
- [15] R.A. Toth, J. Opt. Soc. Am. B10 (1993) 1526-1544.
- [16] R.A. Toth, J. Mol. Spectrosc. 166 (1994) 184–203.
- [17] R.A. Toth, C. Camy-Peyret, J.M. Flaud, J. Mol. Spectrosc. 67 (1977) 185–205.
- [18] J.P. Chevillard, J.Y. Mandin, J.M. Flaud, C. Camy-Peyret, Can. J. Phys. 63 (1985) 1112–1127.
- [19] R.A. Toth, Appl. Opt. 33 (1994) 4868-4879.
- [20] J.P. Chevillard, J.Y. Mandin, J.M. Flaud, C. Camy-Peyret, Can. J. Phys. 65 (1987) 777–789.
- [21] J.P. Chevillard, J.Y. Mandin, J.M. Flaud, C. Camy-Peyret, JQSRT 36 (1986) 395–399.
- [22] A.S. Pine, M.J. Coulombe, C. Camy-Peyret, J.M. Flaud, J. Phys. Chem. Ref. Data 12 (1983) 413–465.
- [23] J.M. Flaud, C. Camy-Peyret, J.P. Maillard, Mol. Phys. 32 (1976) 499–521.
- [24] L. Wallace, P.F. Bernath, W. Livingstone, K. Hinkle, J. Busler, B. Guo, K. Zang, Science 268 (1995) 1155–1158.
- [25] O.L. Polyansky, N.F. Zobov, J. Tennyson, J.A. Lotoski, P.F. Bernath, J. Mol. Spetrosc. 184 (1997) 35–50.
- [26] M.P. Esplin, R.B. Wattson, M.L. Hoke, L.S. Rothman, JQSRT 60 (1998) 711–739.
- [27] R. Lanquetin, L.H. Coudert, C. Camy-Peyret, J. Mol. Spectrosc. 195 (1999) 54–67.
- [28] S.N. Mikhailenko, Vl.G. Tyuterev, V.I. Starikov, K.A. Albert, B.P. Winnewisser, M. Winnewisser, G. Mellau, C. Camy-Peyret, J.M. Flaud, J.W. Brault, J. Mol. Spectrosc. 213 (2002) 91–121.
- [29] Vl.G. Tyuterev, J. Mol. Spectrosc. 151 (1992) 97-129.
- [30] VI.G. Tyuterev, V.I. Starikov, S.A. Tashkun, S.N. Mikhailenko, J. Mol. Spectrosc. 170 (1995) 38–58.
- [31] V.I. Starikov, Vl.G. Tyuterev, Intramolecular Rovibrational Interactions and Theoretical Methods in the Spectroscopy of Non-rigid Molecules, Nauka, Moscow, 1997, pp. 1–223 [in Russian].

- [32] J.K.G. Watson, J. Chem. Phys. 46 (1967) 1936-1949.
- [33] H. Partridge, D.W. Schwenke, J. Chem. Phys. 106 (1997) 4618– 4639.
- [34] M. Birk, M. Winnewisser, E.A. Cohen, J. Mol. Spectrosc. 136 (1989) 402–445.
- [35] P.S. Ormsby, K. Narahari Rao, M. Winnewisser, B.P. Winnewisser, O.V. Naumenko, A.D. Bykov, L.N. Sinitsa, J. Mol. Spectrosc. 158 (1993) 109–130.
- [36] S.N. Mikhailenko, Vl.G. Tyuterev, K.A. Keppler, B.P. Winnewisser, M. Winnewisser, G. Mellau, S. Klee, K.N. Rao, J. Mol. Spectrosc. 184 (1997) 330–349.
- [37] W. Quapp, M. Hirsch, G.Ch. Mellau, S. Klee, M. Winnewisser, A. Maki, J. Mol. Spectrosc. 195 (1999) 284–298.
- [38] A. Maki, G.Ch. Mellau, S. Klee, M. Winnewisser, W. Quapp, J. Mol. Spectrosc. 202 (2000) 67–82.
- [39] L.S. Rothman, C.P. Rinsland, A. Goldman, S.T. Massie, D.P. Edwards, J.M. Flaud, A. Perrin, C. Camy-Peyret, V. Dana, J.Y. Mandin, J. Schroeder, A. McCann, R.R. Gamache, R.B. Wattson, K. Yoshino, K.V. Chance, K.W. Jucks, L.R. Brown, V. Nemtchinov, P. Varanasi, JQSRT 60 (1998) 665–710.
- [40] S.A. Tashkun, S.N. Mikhailenko, Vl.G. Tyuterev, J.L. Teffo, I. Perevalov, in preparation.